1	Carbonatitic magma fractionation and contamination generate rare earth
2	element enrichment and mineralization in the Maoniuping giant REE deposit,
3	SW China
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22 ABSTRACT

23 Carbonatite intrusions host the world's most important light rare earth element 24 (LREE) deposits, and their formation generally requires extraordinary fertile sources, magmatic evolution, and hydrothermal events. However, carbonatitic magma 25 evolution, particularly the role of fractional crystallization and contamination from 26 silicate rocks in REE enrichment, remains enigmatic. The Maoniuping world-class 27 28 REE deposit in southwestern China, is an ideal target to decipher magmatic evolution and related REE enrichment as it shows continuous textual evolution from medium-29 to coarse-grained calcite carbonatite (carbonatite I) at depth, to progressively 30 31 pegmatoidal calcite carbonatite (carbonatite II) at shallow levels. In both types of calcite carbonatites, four generations of calcite can be classified according to 32 petrographic and geochemical characteristics. Early-crystalizing calcite (Cal-I and 33 34 Cal-II) are found in carbonatite I and exhibit equigranular and a polygonal mosaic textures, while late calcites (Cal-III and Cal-IV) in carbonatite II are large-size 35 oikocrysts (>0.5 mm in length) with strain-induced undulatory extinction and bent 36 twinning lamellae. All these generations of calcite yield similar, near-chondritic, Y/Ho 37 38 ratios (26.6–28.1) and are inferred to be of magmatic origin. Remarkably, gradual 39 enrichment of MgO, FeO and MnO from Cal-I to Cal-IV is coupled with a significant 40 increase in REE contents (~800 to 2000 ppm), with LREE-rich and gentle-to-steep 41 chondrite-normalized REE patterns ((La/Yb)_N = 3.1-26.8 and (La/Sm)_N = 0.9-3.9, respectively). Such significant REE enrichment is ascribed to protracted magma 42

43	fractional crystallization with initial low degree of fractional crystallization (fraction
44	of melt remining (F) = ~0.95) evolving to late stage (F = $0.5-0.6$) by formation of
45	abundant calcite cumulates. Differential LREE and HREE behavior during magma
46	evolution largely depend on separation of phlogopite, amphibole, and clinopyroxene
47	from the carbonatitic melt, which is indicated by progressively elevated $(La/Yb)_N$
48	ratios ranging from 3.1 to 26.8. The four generations of calcite have significantly
49	different C and Sr isotopic compositions with $\delta^{13}C_{V-PDB}$ decreasing from -3.28 to -
50	9.97 ‰ and 87 Sr/ 86 Sr increasing from 0.70613 to 0.70670. According to spatial
51	relations and petrographic observations, the relative enrichment of $\delta^{13}C$ and depletion
52	in ⁸⁷ Sr/ ⁸⁶ Sr ratios of Cal-I and Cal-II show primary isotopic characteristics inherited
53	from initial carbonatitic magma. By contrast, the variable Sr and C isotopic
54	compositions of Cal-III and Cal-IV are interpreted as the results of contamination by
55	components derived from silicate wall rocks and loss of CO2 by decarbonation
56	reactions. To model such contamination processes, Raleigh volatilization and Monte
57	Carlo simulation have been invoked and the model results reveal that carbonatitic
58	melt-wall rock interaction requires 40% radiogenic Sr contamination from silicate
59	rocks and 35% CO ₂ degassing from carbonatitic melt. Moreover, positive correlations
60	between decreasing $\delta^{13}C$ values and increasing REE contents, together with
61	bastnäsite-(Ce) precipitation, indicate further REE accumulation during the
62	contamination processes. In summary, alongside REE-rich magma sources, the extent
63	of fractional crystallization and contamination during carbonatitic magma evolution

are inferred to be important mechanisms in terms of REE enrichment andmineralization in carbonatite-related REE deposits worldwide.

Keywords: calcite carbonatite; fractional crystallization of carbonatitic magma;
wall-rock contamination; Sr-C isotopic decoupling; REE enrichment and
mineralization

69 **INTRODUCTION**

70 Rare earth elements (REEs) are essential raw materials for aerospace, permanent magnets, electric vehicles and renewable energy generation as part of the group of 71 72 so-called critical metals (Chakhmouradian & Wall, 2012, Linnen et al., 2014; Wall &Pell, 2020). Carbonatite-related REE deposits are the world's most important source 73 of LREEs and host approximately 51.4% of global rare earth oxides (REO) resources 74 (Weng et al., 2015). An increasing body of research indicates that the genesis of 75 carbonatite-related REE deposits is not the product of a simple one-stage process, but 76 involved a sequence of extraordinary REE enrichment in metasomatic mantle sources 77 followed by magmatic to hydrothermal evolution (e.g., Bayan Obo, China, Smith & 78 79 Henderson, 2000; Smith et al., 2007, 2015; Yang et al., 2019; Wei et al., 2022; 80 Mountain Pass, USA, Castor, 2008; Poletti et al., 2016; Maoniuping, China, Hou et al., 2006, 2015; Liu et al., 2019a, b; Zheng & Liu, 2019; Bear Lodge, USA, Moore et al., 81 82 2015; Andersen et al., 2016); or external hydrothermal alteration and/or supergene 83 weathering events (e.g., Mount Weld, Australia, Jaireth et al., 2014; Zandkopsdrift, 84 South Africa, Venter et al., 2010; Harper et al., 2015). Thus, deciphering the formation

85 processes of carbonatite-related REE deposits is challenging, which requires integrated studies of geochemistry of the principle host-rocks and, textural and 86 87 compositional variations of their constituent minerals (Chakhmouradian & Zaitsev, 88 2012; Goodenough et al., 2021; Anenburg et al., 2022). Rare earth element behavior in carbonatite magma sources, magma evolution 89 90 and hydrothermal modification is relatively poorly constrained by limited numbers of experiments, modeling studies and relatively few natural case studies (Veksler et al, 91 1998; Bühn, 2008; Chakhmouradian et al., 2015, 2016a, b; Chebotarev et al., 2019, 92 93 2022; Yang et al., 2019; Anenburg et al., 2020b; Nably et al., 2020; Walter et al., 2022, 2023). The study of these processes is hindered by the fact that carbonatites are easily 94 modified by hydrothermal reworking, structural events or multiple overlapping 95 96 processes, resulting in difficulty in recognizing the effects of magmatic processes (Chakhmouradian et al., 2016b; Yaxley et al., 2022). Magmatic evolution processes 97 responsible for REE enrichment are commonly complicated and in many systems may 98 99 be unique due to extensive crystallization and contamination in carbonatite systems 100 (Bühn et al., 2001; Chen & Simonetti, 2013; Giebel et al., 2019a, b; Anenburg et al., 101 2020a, b; Fomina & Kozlov, 2021; Bouabdellan et al., 2022; Chmyz et al., 2022; 102 Walter et al., 2022, 2023). For instance, REE accumulation in the world's largest REE 103 deposit at Bayan Obo has been inferred to be controlled by an unusual magmatic 104 evolution history from ferroan-, followed by magnesian- and finally calcio-carbonatite 105 varieties (Yang et al., 2011, 2019). However, this evolutionary sequence is not

106	representative for most known carbonatite intrusions which typically evolved from
107	calcio-, through magnesian- to ferroan-carbonatite series (Anenburg et al., 2022).
108	Over 60% carbonatites worldwide are predominantly calcite and most economic REE
109	mineralization is associated with Ca-rich carbonatites (Woolley & Kjarsgaard, 2008;
110	Chakhmouradian & Zaitsev, 2012; Smith et al., 2016). However, recent studies
111	indicate high REE concentrations in residual ferroan dolomite carbonatite melt
112	(Anenburg et al., 2022). Given that REE redistribution in the latter processes is results
113	of back reaction with precursor alkali-REE-carbonates (e.g., burbankite and
114	carbocernaite) by hydrothermal fluids expelled from the brine-melt stage (Prokopyev
115	et al., 2016; Anenburg et al., 2022), primary REE enrichment in calciocarbonatitic
116	melt may be a prerequisite. Thus, detailed studies on the evolutionary history of
117	calcio-carbonatitic melt and its relationship to REE enrichment are necessary. In
118	addition, contamination via reaction with silicate rocks during carbonatitic magma
119	ascent is considered to influence REE mineralization by a series of interactions
120	between carbonatite melt and silicate rocks (Anenburg & Mavrogenes, 2018; Giebel
121	et al., 2019a, b; Walter et al., 2021). Generally, this process cannot be easily identified
122	by bulk analyses of radiogenic Sr, Nd and Pb isotopic compositions because these
123	element concentrations in carbonatites are enriched by several orders of magnitude
124	compared to those in associated silicate counterparts (Bell &Tilton, 2002). Equally
125	contamination is limited by the rapid ascent of carbonatite magma, owing to their low
126	density and extremely low viscosity (Jones et al., 2013; Kono et al., 2014).

127 Identification of contamination processes and their relationship to REE enrichment during carbonatitic magma ascent from integrated petrographic studies and isotopic 128 129 measurement of Sr, Nd and C could be critical in determining genetic models for REE 130 mineralization systems, but such studies are very sparse (e.g., Giebel et al., 2019b; 131 Fomina & Kozlov, 2021). 132 Calcite is the crucial magmatic carbonate mineral that accounts for the bulk of the main constituents in most carbonatites. Owing to the fact REE can be incorporated 133 into calcite by coupled substitution with Na, calcite is an important REE carrier in 134 135 carbonatite and it's textural and REE budget can serve as a monitor for carbonatitic magma evolution (Chakhmouradian & Zaitsev, 2012; Chakhmouradian et al., 2016b). 136 On the one hand, textural and compositional signatures of calcite, especially REE 137 138 contents and patterns, can be used to trace evolution of carbonatite magma (Chen & Simonetti, 2013; Milani et al., 2017; Ying et al., 2020; Zeng et al., 2022), and ratios of 139 Ce/Ce^{*} and Eu/Eu^{*} are able to indicate their redox state (Chakhmouradian et al., 140 2016a, b; Chmyz et al., 2022). On the other hand, the in-situ Sr and C isotopic 141 142 composition of calcite, coupled with mineralogy and mineral chemistry of other proxy 143 minerals (e.g., apatite, mica, amphibole and clinopyroxene) are beneficial for 144 investigating contamination by silicate rocks, and their roles in REE enrichment 145 during carbonatitic magma evolution (e.g., Reguir et al., 2009, 2012: Chakhmouradian et al., 2017; Anenburg & Mavrogenes, 2018; Giebel et al., 2019a, b; 146 147 Walter et al., 2022, 2023).

148 Carbonatites from the Maoniuping word-class REE deposit (southwestern China) have been investigated previously to determine the origin of carbonatite, fluid 149 150 evolution, and REE mineralization (Yuan et al., 1995; Wang et al., 2001; Xu et al., 151 2003; Hou et al., 2006, 2015; Xie et al., 2009; Liu et al., 2015, 2019a, b; Zheng & Liu, 2019; Weng et al., 2021, 2022a, b). Here, we provide new findings on the integrated 152 153 petrology, mineralogy, mineral chemistry, and in situ Sr and C isotopic composition of calcite and silicate minerals in different types of calcite carbonatites from a new 154 1100m drill core. These geochemical features and field observations outline a 155 conceptual model for REE enrichment and mineralization based on magma 156 fractionation and wall-rock contamination. Furthermore, through crystallization 157 modeling, calculations of Raleigh volatilization, and Monte Carlo simulations of 158 159 contamination, we infer in detail the processes of fractional crystallization and contamination-induced interaction between carbonatitic melt and silicate rocks and 160 assess how the both key processes identified at Maoniuping may have influenced 161 REE-rich carbonatite generation globally. 162

163

164 **GEOLOGICAL SETTING**

The 270-km-long Mianning-Dechang (MD) belt in eastern Tibetan hosts a total resource of more than 3.2 Mt of REOs and comprise the Maoniuping, Dalucao, Muluozhai and Lizhuang REE deposits (Fig. 1a and b, Yin & Harrison, 2000; Shu & Liu, 2019; Shu et al., 2020; Zhang et al., 2022). This economic metallogenic belt 169 formed during the late stage of Indian-Asian continental collision (40-24 Ma, Hou & Cook, 2009; Liu & Hou, 2017) and is related to regionally structural activities 170 involving N-S-trending strike-slip movement, ductile shear and transpressional 171 172 deformation (Fig. 1b; Yin & Harrison, 2000; Deng et al., 2014, 2021; Zeng & Liu, 2022). Carbonatites in the MD belt occur as stocks, dykes, and sills associated with 173 174 nordmarkite that is a variety of quartz-bearing alkali feldspar syenite (Yuan et al., 1995). Petrological studies suggest that carbonatites and nordmarkite originate by 175 liquid immiscibility of CO₂-rich silicate melt (Hou et al., 2006), and their sources are 176 177 related to the sub-continental lithospheric mantle (SCLM) that had been previously metasomatized by high-fluxing elements and CO2-rich fluids from subducted marine 178 sediments (Hou et al., 2015). 179 180 The Maoniuping carbonatite-related REE deposit is located at the north part of the MD belt and close to the Yalongjiang strike-slip fault (Fig. 1b). The currently 181 explored resources at Maoniuping comprise more than 3.17 Mt REO at an average 182 grade of 2.95 wt.%, and accessory products of 0.6 Mt Pb (average grade = 0.44 wt.%), 183 19.67 Mt barvte (average grade = 8.03 wt.%), 13.24 Mt fluorite (average grade = 6.70184 wt.%), and 0.83 Mt celestine (average grade = 1.03 wt.%) (Yuan et al., 1995). The 185 186 host rock geology at Maoniuping includes five main lithological units (Fig. 1c, Yuan 187 et al., 1995; Liu and Hou, 2017): (1) a series of N-S-trending granitic plutons, 188 composed of pink alkalic granite, light gray alkalic granite, and graphic granite; (2) an

189 approximately 1100-m-thick sequence of metamorphic Devonian-Permian clastic

190 rocks, limestone, and flood basalts intruded by Mesozoic granite; (3) a Paleozoic rhyolite dated at 600-750 Ma; (4) a 700-m-thick coal-bearing Triassic sedimentary 191 sequence overlying the metamorphic basement; and (5) a NE-SW-trending 192 193 carbonatite-nordmarkite complex, 1400 m long and 260-350 m wide. The distribution and orientation of all the units is controlled by the Haha strike-slip fault (Fig. 1c). At 194 195 Maoniuping, the carbonatite intruded the cogenetic nordmarkite or Mesozoic alkalic granite as stocks, dykes, and sills. Sensitive high-resolution ion microprobe zircon 196 U-Pb age of 26.6 \pm 0.3 Ma (Ling et al., 2016) and arfvedsonite/biotite ⁴⁰Ar-³⁹Ar age 197 198 of carbonatite of 27.6 ± 2.0 Ma (Liu & Hou, 2017), have been reported for the 199 nordmarkite and carbonatite, respectively. These ages correspond within error to the SIMS (secondary ion mass spectrometry) U-Th-Pb age of 26.8 ± 0.7 Ma for 200 201 bastnäsite-(Ce) from carbonatite (Weng et al., 2022a), indicating contemporaneous carbonatite genesis and REE mineralization around 26 Ma at Maoniuping. Previous 202 studies show that most carbonatites are located in the north of the deposit 203 (Guangtoushan section) and are characterized by coarse-grained calcite (50–90 vol.%) 204 205 with a subordinary mineral paragenesis of aegirine/aegirine-augite, arfvedsonite, 206 fluorite, baryte, and minor bastnäsite-(Ce) in outcrop (Zheng & Liu, 2019). 207 REE mineralization at Maoniuping is hosted inside carbonatites or 208 carbonatite-derived veins that crosscut nordmarkite as REE mineralized ore veins that 209 show "S"-like outline (Fig. 1c; Yuan et al., 1995; Liu & Hou, 2017). Such REE 210 mineralized ore veins are well developed in the Dagudao section, which is currently

211 the main open pit, and are composed of low-grade stockworks of multiple veinlets in 212 the lower part of the orebody, grading upwards into progressively thicker veins (Liu et 213 al., 2019a). In the top unit, ore veins can be categorized into four different types based 214 on mineral assemblages and texture (Liu et al., 2019a). The dominant types of ore veins are calcite veins that are typically zoned and comprise K-feldspar (microcline 215 216 and orthoclase), micas (phlogopite), amphibole (arfvedsonite), pyroxene (aegirine to aegirine-augite), calcite, fluorite, baryte, and bastnäsite-(Ce) (Zheng & Liu, 2019). 217 Previous studies propose that carbonatitic fluids are responsible for the main REE 218 219 mineralization as most fluid inclusions in bastnäsite-(Ce) are characterized by 220 meteoric water-rich, carbonate-bearing fluids and have homogenization temperatures below 240 °C (Xie et al., 2015; Guo & Liu, 2019; Zheng & Liu, 2019). However, 221 222 recent research has argued that a portion of bastnäsite-(Ce) in carbonatites is of magmatic origin based on their contained melt inclusions and discriminably 223 magmatic-like low U contents (<70 ppm), low (La/Yb)_n ratios (<~30000) and Eu 224 deficiency ($\delta Eu = 0.67 - 0.98$) compared to hydrothermal bastnäsite-(Ce) (Weng et al., 225 226 2022a)

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228 SAMPLING AND ANALYTICAL METHODS

A total of 44 samples were collected from the first 1100 m of the HZKS drill core (Figs. 1c and 2). Petrographic studies were conduct using a Zeiss Axioskop 40 optical microscope and backscattered electron (BSE) imaging were acquired using FEI Quanta 450 SEM with an accelerating voltage of 20 kV and a beam current of 30

233 nA. Representative samples were selected for optical microscope

234 cathodoluminescence (OP-CL) characterization using a Leica DM2700P microscope

- 235 coupled with a Cambridge Image Technology LTD (CITL) MK 5-2 system at 12 kV
- accelerating voltage and 350 mA current in China University of Geosciences (CUG),
- 237 Wuhan, China.

Whole-rock chemical analyses of major and trace elements were carried out at 238 ALS Chemex (Guangzhou) Co., Ltd., China, by X-ray fluorescence (XRF) 239 spectrometry (Fluorine mode) on lithium tetraborate-lithium metaborate flux fused 240 glass disks. Certified reference material (CRM) TDB-1 served as external standard for 241 Cl and multiple reference materials (GBW07241, GBW07237, GBW07211) (Table S1) 242 243 served as external standards for major and trace elements. Analytical precision was better than $\pm 2\%$ for major element oxides (SiO₂, Al₂O₃, Fe₂O₃^T, CaO, Na₂O, K₂O, 244 MgO, MnO, TiO₂). To obtain Fe^{2+} and Fe^{3+} contents, more than 200 mg of powder, 245 separated from the same disks, was digested with a mixture of H₂SO₄-HF, and was 246 then titrated in a beaker using $K_2Cr_2O_7$ to obtain Fe^{2+} concentration. For trace 247 248 elements, powders of each sample were accurately weighed to 0.100 g and then mixed 249 with $Li_2B_4O_7$ or $LiBO_2$ (specifically for analyses of REE concentrations) in a Savillex 250 Teflon beaker. The mixture was heated for 24 h at 80 °C and then dissolved in an acid 251 mixture of 1.5 ml HNO₃, 1.5 ml HF, and 0.5 ml HClO₄ for digestion in a high-temperature oven at 180 °C for 72 h. After cooling, the sample solution was 252

concentrations were analyzed using an Agilent 7700x quadrupole inductively coupled 254 plasma mass spectrometer (ICP-MS). Multiple reference materials (OGGeo08, 255 256 MRGeo08, OREAS-120, OREAE-45e, GBM908-10, STSD-1) (Table S1) served as external standards, and analytical precision was better than 10%. 257 258 The carbon and oxygen isotopic composition of calcite from carbonatites was determined at CAGS. The selected samples were washed with distilled water and 259 dried at room temperature. Samples were then crushed to less than 200 mesh in an 260 261 agate mortar. Each powder was treated with 100% anhydrous H₃PO₄ at 50 °C to generate CO₂ that was subsequently purified, and then C and O isotopic compositions 262 were measured using a Finnigan MAT-251 isotope ratio mass spectrometer. The 263 analytical uncertainty is better than 0.15% and 0.020% (2 σ) for C and O isotope 264 ratios, respectively. Isotopic compositions reported here are normalized to the Vienna 265 Pee Dee Belemnite standard (V-PDB) for δ^{13} C and Vienna Standard Mean Ocean 266 Water (V-SMOW) for δ^{18} O. 267

transferred to a plastic bottle and diluted with 1% HNO₃ to 50 ml. The trace element

253

The major element compositions of silicate minerals and calcite were determined
using a JXA-iHP200F electron microprobe at the Institute of Mineral Resources,
CAGS. Accelerating voltage was 15 kV using a beam current of 10 nA (for carbonates)
and 20 nA (for other minerals). The beam diameter was 5 µm for silicates and 10 µm
for calcite. All elements were measured by peak mode and the peak counting times for
F was 40s; for all the other elements it was set to 20 s on peak and 10 s on background.

Natural and synthetic standards were used for calibration (Table S1) and raw data were corrected using ZAF procedures. The elements Al, Ba, F, Fe, K, Mg, Mn, Na, Si, and Ti were mapped for phlogopite using the same instrument at an accelerating voltage of 15 kV and a beam current of 10 nA. A total of 602×476 points were gathered for each element map with dwell time of 50 ms for each point and a probe diameter of 5 µm.

The in-situ trace element analyses of minerals were conducted using a 280 RESOlution 193 nm laser ablation (LA) system coupled to a Thermo iCAP-RQ 281 282 ICP-MS in the State Key Laboratory of Geological Process and Mineral Resources (GPMR), CUG. Laser ablation spots were located on the same crystals analyzed by 283 EPMA, with energy density of $2.8-3.5 \text{ J/cm}^2$ at a repetition rate of 8-10 Hz and a 284 285 beam diameter of 33-50 µm depending on the size of mineral grains. Multiple reference materials (NIST 610, NIST 612, BIR-1G, BCR-2G and BHVO-2G, Jochum 286 et al., 2005) served as external standards (Table S1), which together with Ca from 287 EPMA, were used for concentration calibration (Liu et al., 2008). The precision and 288 accuracy are better than 5% for REEs and 10% for all other trace elements (1 r RSD; 289 290 Chen et al., 2011). Data were processed using an Excel-based software 291 ICPMSDataCal 12.2 via off-line selection and integration of background and 292 analytical signals, and time drift correction and quantitative calibration (Liu et al., 293 2008; Lin et al., 2016). Time resolved mass spectra were checked to identify and 294 exclude analyses affected by mineral micro-inclusions.

295	The in-situ carbon isotope composition of calcites from selected carbonatite
296	intrusions was determined on a RESOlution 193 nm laser ablation system coupled to
297	a Nu Plasma II multicollector (MC)-ICP-MS at GPMR, CUG. $^{13}\mathrm{C}$ and $^{12}\mathrm{C}$ were
298	measured simultaneously in Faraday cups H3 and L4, and the ²⁵ Mg ²⁺ ion signal was
299	acquired in the central cup (Lu et al., 2022). All measurements were performed in low
300	mass resolution mode of ~400, whereas medium mass resolution (~4000) was adopted
301	to resolve spectral interferences (Lu et al., 2022). The samples were ablated in a
302	helium atmosphere (0.3–0.4 L min ⁻¹) mixed with 4 mL min ⁻¹ N ₂ , which was then
303	combined with argon gas (~ 0.6 Lmin^{-1}) before entering the plasma torch (Chen et al.,
304	2017). Analyses were performed using a laser beam diameter of 193 μ m with energy
305	density of 3 J/cm ² and repetition rate of 12 Hz. Each analysis consists of 30 s of
306	background measurement, 40 s of ablation, and 50 s of washout. A standard-sample
307	bracketing (SSB) method was adopted to correct for instrumental mass bias with the
308	external standard SXD8 calcite (Table S1). Every 5 analyses were bracketed by an
309	analysis of an obtained Oka153 calcite (Chen et al., 2017) as the monitor standard to
310	determine the accuracy and precision. The analytical $\delta^{13}C$ values were calibrated
311	using Iso-Compass software (Zhang et al., 2020) for spectral interferences from
312	doubly charged ions (i.e., ${}^{24}Mg^{2+}$ for ${}^{12}C^+$ and ${}^{26}Mg^{2+}$ for ${}^{13}C^+$) and analytical
313	uncertainty is better than 0.25‰ (1 σ) for C isotopes based on repeated measurement
314	of the SXD8 calcites (Lu et al., 2022).

315 In-situ Sr isotopic measurements of calcites were conducted on the same LA-(MC)-ICP-MS instrument configuration at GPMR. Laser ablation spots for Sr 316 isotopic measurement were carried out on the same calcite grains analyzed for in-situ 317 318 C isotope compositions. All analyses were conducted with a laser beam diameter of 75 μ m with an energy density of 3 J/cm² and repetition rate of 10 Hz. Every 4-6 319 320 sample analyses were bracketed by analysis of an in-house coral (Qingdao) as external standard to check the analytical reliability and stability. The average ⁸⁷Sr/⁸⁶Sr 321 ratio obtained for the coral standard was 0.70918 ± 0.00004 (25, n = 65, Table S1), 322 corresponding to the values of 0.70910 ± 0.00002 to 0.70923 ± 0.00002 reported by 323 ID-TIMS (Bizzarro et al., 2003; Chen et al., 2018). Double-charged ions of ¹⁶⁷Er²⁺, 324 ¹⁷¹Yb²⁺, ¹⁷³Yb²⁺ were monitored based on the method of Ramos et al. (2004). To 325 exclude the interferences of Kr and Rb, the interferences of ⁸⁴Kr on ⁸⁴Sr, ⁸⁶Kr on ⁸⁶Sr, 326 and ⁸⁵Rb on ⁸⁷Sr were corrected based on the isotopic ratios of 83 Kr/ 84 Kr = 0.20175, 327 83 Kr/ 86 Kr = 0.66474, and 85 Rb/ 87 Rb = 2.5926, respectively. The 87 Sr/ 86 Sr ratios were 328 calculated and calibrated from the interference-corrected ⁸⁶Sr/⁸⁸Sr using the 329 exponential law method described by Yang et al. (2014). 330 331

332 PETROGRAPHY AND MINERALOGY

A simplified log that illustrates the lithological variations in the HZKS-1 drill core is shown in Figure 2. A rock paragenesis of alkaline granite, nordmarkite, and and mineralogical characteristics are shown in Figures 3–6.

337 Alkaline granite

338	The alkaline granite is found throughout the drill core, down to 950 m (Fig. 2)
339	and is gray in color and mostly medium- to fine grained with equigranular texture. It
340	is characterized by the presence of alkali feldspar and quartz with minor biotite,
341	zircon, and apatite (Fig. 3c). The principal minerals in fresh alkaline granite are
342	microcline (~30 vol.%), perthite (~20 vol.%), orthoclase (~15 vol.%), albite (~10
343	vol.%) and quartz (20–25 vol.%). Tabular microcline crystals are 0.5 to 2 mm in size
344	and share planar contacts with lath-shaped perthite. Albite grains are 50–100 μm in
345	size and primary quartz occurs as subhedral to anhedral crystals interstitial to perthite
346	and microcline (Fig. 3c).
347	In the top ~ 100 m of the drill core, alkaline granite is weathered, appears
348	leached, and has a loose or porous texture (Fig. 3e). In proximity to carbonatite II, the
349	alkaline granite exhibits carbonation and REE mineralization, which is distinct from

alkaline granite exhibits carbonation and REE mineralization, which is distinct from
fresh granite. The altered granite is reddish in color; contains platy fluorite, baryte,
and bastnäsite-(Ce); and hosts comb-like aggregates of amphibole and clinopyroxene
at the contact with carbonatite (Fig. 3e).

353

355 Nordmarkite spatially occupies in the drill core at 190–550, 890–908, 924–933, and 845–1100 m. In hand specimen, fresh nordmarkite is gray- to white (Fig. 3a) and 356 357 consists of alkali feldspar, quartz, pyroxene, and phlogopite with accessory zircon. apatite, magnetite, pyrite, and magnetite. Alkali feldspar accounts for over 70 vol.% 358 of the rock and can be subdivided into microcline (30 vol.%), orthoclase (30 vol.%) 359 and albite (10 vol.%) (Fig. S1a and b). Albite occurs along the cleavage of orthoclase 360 grains as irregularly perthitic crystals (Fig. S1a). In some places, feldspar crystals are 361 362 partly altered to sericite, which increases appreciably with proximity to the contact 363 with carbonatite II (Fig. S1c and e). Quartz occurs interstitially to microcline and 364 orthoclase (Fig. S1d). Pyroxene occurs as skeletal crystals or aggregates that usually 365 contain elongated or isometric apatite, euhedral titanite and bastnäsite-(Ce) (Fig. S1b, 366 d and f).

367

368 Calcite carbonatite

369 Based on textural characteristics and mineral assemblages, two varieties of
370 carbonatites are recognized, including medium- to coarse-grained carbonatite
371 (carbonatite I, Fig. 3g–i) and super coarse-grained carbonatite (carbonatite II, Fig. 3j–
372 k). Generally, carbonatite I is found at greater depths and gradually coarsens upward
373 into carbonatite II that contains numerous centimeter-scale grain size intervals and

375 can be distinguished by variable proportions of silicate minerals (e.g., phlogopite,

376 riebeckite-arfvedsonite, aegirine-augite), fluorite, and baryte. Additional details that

- 377 distinguish carbonatite I and II are presented below:
- 378

The medium- to coarse-grained calcite carbonatite (carbonatite I)

379 The carbonatite I mainly occurs from 953 to 730 m in the deep drill core (Fig. 2) 380 and intrudes nordmarkite and alkaline granite as veins, dykes and plugs (Figs. 2 and 3d, f and j) less than 1 m in width. This variety exhibits gray to white color, 381 equigranular texture and is comprised of overwhelmingly fine-grained calcite (<500 382 383 µm in diameter and 85 to 90 vol.%), ferromagnesian silicates (phlogopite and riebeckite, 5–10 vol.%), and a handful of magnetite and molybdenite grains (total <1 384 vol.%) (Fig. 4a-d). Euhedral- to subhedral calcite crystals are equant, and oval, or 385 lobate in shape and 100 to 500 µm in size (Fig. 4a). The fine grains, which are 386 387 transparent, show colorful extinction and weak twinning in cross-polarized light (Fig. 4c) and mid-gray color in BSE images (Fig. 4e). Under OP-CL imaging, calcite 388 389 displays bright orange to blood-red grains that have slight core to rim zonation (Fig. 390 4i). The boundary between calcite crystals is usually planar, which, combined with the 391 lack of mechanical twinning, shows that they have not undergone textural 392 re-equilibration. Phlogopite in carbonatite I is 0.2 to 0.5 mm in size and occurs 393 interstitially to calcite with the same planar contacts (Fig. 4b). Most phlogopite grains 394 display dark-green color in plane-polarized light and are disseminated in carbonatite I

395	(Figs. 3g-h and 4b). Riebeckite is euhedral prismatic or needle-like crystals enclosed
396	in phlogopite as compact inclusions (<50 μm across) (Fig. 4b and f). Accessory
397	minerals include magnetite and molybdenite that form aggregates or clusters in local
398	parts of carbonatite I (Fig. 4d).
399	Characteristics of carbonatite I can vary over short distances, on the order of tens
400	of meters (Fig. 4g and h). These carbonatites are pink and show calcite oikocrysts
401	containing tiny bastnäsite-(Ce) inclusions (Fig. 4g). In BSE images, the bastnäsite-(Ce)
402	inclusions (commonly <10 μ m in size) are orientated along calcite cleavages or occur
403	as broken trails of elongated, spindle- or pill-like crystals (Fig. 4h), whose abundance
404	ranges significantly from absent to as much as $1-2$ vol.%.
405	The super coarse-grained calcite carbonatite (carbonatite II)

The carbonatite II occupies sections at 800-895 m, 730-760 m, 650-680 m, and 406 280-300 m throughout the HZKS-1 drill core (Figs. 2 and 5). These samples 407 commonly exhibit inequigranular and sometimes pegmatoidal texture (Figs. 5d), and 408 are typically represented by large calcite crystals range from commonly 2-5 409 410 millimeters (Fig. 5) to locally several centimeters in size (Fig. 3k). In some parts, 411 closely anchimonomineralic calcite carbonatite can be found (Figs. 3j). Statistically, 412 most carbonatite II is comprised of dominant calcite (60-70 vol.%), fluorite (~10 413 vol.%), and baryte (~10 vol.%) grains accompanied by variable amounts of 414 phlogopite (~5 vol.%), magnesioarfvedsonite (~5 vol.%), aegirine (~5 vol.%), and minor microcline, quartz, pyrite, and galena (Fig. 5a-d, h-k). Bastnäsite-(Ce) is 415

416	widespread and volumetrically accounts for over 10% of carbonatite II in some cases
417	(Fig. 5b, j). Calcite in carbonatite II is subhedral to anhedral, elongated or anhedral
418	crystals (Fig. 5d) and shows complex optical characteristics with subtle undulatory
419	extinction, thick twin lamellae in cross-polarized light (Fig. 5e), and flame-like cores
420	fading to a dark orange rim under OP-CL images (Fig. 5k). Locally, misoriented
421	microfractures show up and obliquely traverse the cleavage of calcite crystals, whose
422	grain boundaries become serrated and bulging (Fig. 5f). These textual characteristics
423	suggest a response to stress and ductile deformation of carbonatite II.
424	Fluorite in carbonatite II occurs as subhedral, elliptical to elongated aggregates
425	disseminated in the calcite-dominant matrix (Figs. 3k and 5h-k), and is dark blue in
426	CL images (Fig. 5k). Boundaries between fluorite and calcite are planar contacts
427	without crosscutting or intersection, representing equilibrium crystallization. Baryte
428	exhibits patchy texture and commonly occurs as anhedral aggregates along calcite
429	fractures (Fig. 5i-j), whereas quartz, which is absent in carbonatite I, is 0.5 to 1 mm in
430	size and occurs interstitially to calcite with dark gray color in BSE images (Fig. 5i).
431	Arfvedsonite (or riebeckite) shows compact, comb-like, euhedral configuration (Fig.
432	5c). The REE minerals in carbonatite II are mainly bastnäsite-(Ce), showing large,
433	lath-shaped, and euhedral crystals (>500 μ m) that syntactically intergrow with
434	parisite-(Ce) (Fig. 5j).

436 In general, REE mineralization in the drill core can be mainly found inside carbonatite II and at the contact zone between carbonatite and wall rocks. As 437 438 described above, REE minerals in carbonatite II are dominantly bastnäsite-(Ce) with lesser amounts of parisite-(Ce), both of which exhibit patchy zoning (Fig. 5) and 439 have rosette-like shape alongside fluorite, calcite, and in some places baryte (Fig. 5b 440 441 and i). In contrast, REE mineralization in the contact zone is characterized by the following three key features that are distinct from those in carbonatite II: (1) REE 442 443 minerals: bastnäsite-(Ce) is abundant (over 10 vol.% in some parts) and occurs 444 interstitially to calcite, with which it shares planar contacts (Figs. 3d and 6a); (2) in 445 BSE images most bastnäsite-(Ce) is unzoned and shows lath-shape, euhedral crystal, or random aggregates with 0.2–0.5 cm in length (Fig. 6e–g); (3) a high abundance of 446 447 silicates, including phlogopite, aegirine or aegirine-augite, riebeckite or arfvedsonite and microcline, which are intergrowth with bastnäsite-(Ce) and exhibit core to-rim 448 texture (Fig. 6a-d) 449

450

452

451 **RESULTS**

Whole-rock chemistry

The whole-rock geochemical results are provided in Table 1, normalized REE patterns are illustrated in Figure 7, and major and trace element discriminant diagrams are shown in Figures S2–S3. Of all the analyzed samples, the fresh nordmarkite

456	commonly comprises SiO_2 contents of 68.05–70.09 wt.%, the highest concentrations
457	of Na ₂ O + K ₂ O (means = 11.09 wt.%) (Fig. S2a and b), but relatively low REE
458	contents and low LREE/HREE ratios (208–455 ppm and <87.4, respectively, Figs. 7b,
459	S2e and f). In comparison, nordmarkite samples in contact zones have lower Na ₂ O +
460	K_2O contents (6.50 wt.%), but higher contents of SO_3 (1.27 wt.%), F (2.1 wt.%) and
461	REE (2603 ppm) (Fig. S2e and f). For alkali granite, the fresh alkali granite is
462	characterized by high contents of SiO ₂ (66.87–76.01 wt.%) and (La/Yb) _N ratios
463	(27.0–306), but low contents of CaO, SO ₃ , F, Nb, REE (431–3202 ppm), and ratios of
464	Zr/Hf and Nb/Ta (27.34–33.65 and 35.19–87.39, respectively) (Fig. S2d). In contrast,
465	the altered variety has variable contents of SiO ₂ (38.21-73.40 wt.%), CaO (0.13-
466	16.30 wt.%), higher contents of F (up to 3.10 wt.%), HREE (30.9-1310 ppm), Nb
467	(125–1925 ppm), and low (La/Yb) _N ratios (1.52–67.7; Figs. S4c, f, and i). Moreover,
468	the altered alkali granite is further characterized by tetrad-effect-like REE patterns
469	with pronounced negative Eu anomalies (Fig. 7b), indicating that they were formed
470	from a more fractionated granitic system; the result of a magma-fluid system evolving
471	under open conditions; or the products of a separate magmatic event (Irber, 1999).
472	Both carbonatites I and carbonatite II are consistently calcium-rich (CaO = 41.62
473	to 52.73 wt.%) and exhibit high CaO/ (CaO+MgO+FeO) ratios of 0.96–0.98 (Table 1),
474	and are therefore classified as calcite carbonatites (Fig. S3). Basically, carbonatite I is
475	much lower in SiO ₂ , F, SO ₃ and BaO than carbonatite II (Fig. S4b and e), and has
476	similar low MgO, FeO, MnO, P2O5 (<0.25 wt.%) contents (Fig. S4b, d, e). Their

477	trace-element budget is characterized by low REE contents (3098–3198 ppm), low
478	ratios of LREE/HREE and $(La/Sm)_N$ (12.8–28.9 and 2.75–10, respectively) with
479	slight Eu deficiency ($\delta Eu = Eu_N / (0.5Sm_N + 0.5Gd_N)$: 0.89–0.94), unity δCe ($\delta Ce =$
480	$Ce_N / (0.5La_N + 0.5Pr_N)$: 1.00–1.04), and near-chondritic Y/Ho (22.8–25.8) and Th/U
481	ratios (0.2–3.3). In comparison, the carbonatite samples in contact zones have a wide
482	range of CaO and SiO ₂ contents (Fig. S4a), and higher Al ₂ O ₃ , Na ₂ O, K ₂ O, F and SO ₃
483	contents, but relative depletion in SrO and MnO (Fig. S4b–e).
484	Compared to carbonatite I, carbonatite II has variable SiO ₂ contents (0.15–15.03
105	(1, 0) avident by presence of large properties of phases its alignmentations

wt.%), evident by presence of large proportions of phlogopite, clinopyroxene 485 amphibole, and quartz (Fig. 5c and i). Overall, they contain similarly low P2O5 486 contents (<0.11 wt.%), but are much enriched in BaO, SO₃ and F (up to 13.40 wt.%, 487 5.03 wt.% and 15.70 wt.%, respectively). Due to the large number of bastnäsite-(Ce) 488 grains, carbonatite II hosts high levels of REE (2241–127,686 ppm) and exhibits steep 489 chondrite-normalized REE patterns with over three orders of magnitude variation 490 from La to Lu ((La/Yb)_N = 247-2202; Fig, S4c and h). Moreover, carbonatite II is 491 492 further discriminable in $\delta Ce (0.89-1.33)$ and $\delta Y (\delta Y = Y_N / (0.25 Dy_N + 0.75 Ho_N))$: 493 0.72–1.30) from carbonatite I ($\delta Y = 0.77-0.86$) (Fig, S4i).

495 Mineral composition

496 Calcite

494

497 The chemical compositions of calcite from carbonatite I to II, and contact zones are listed in Tables S2-S3. The formula calculations for calcite are based on one 498 cation with CO2 calculated from stoichiometry. Chemical variations of calcite are 499 500 illustrated in Figures 8 to 10. In general, calcites in all carbonatite varieties contain Ca 501 Mg and small amounts of Fe and with mean formula of 502 (Ca_{0.966}Mg_{0.005}Fe_{0.008}Mn_{0.013}Sr_{0.007})CO₃ All analyzed samples are characterized by discernible Sr (<0.014 in atoms per formula unit (a.p.f.u.)), limited S (<0.002 a.p.f.u.), 503 Y/Ho ratios (23.8 - 34.7)plot within 504 and similar that/ the CHArge-and-RAdius-Controlled (CHARAC) field (Fig. 505 10) where elements of similar charge and radius display extremely coherent behavior and retain their 506 respective chondritic ratio (Bau, 1996). Combining specific textural, major and trace 507 508 characteristics and mineral assemblages, four generations of calcite are identified, namely Calcite-I, Calcite-II, Calcite-III and Calcite-IV. 509 Calcite-I (Cal-I) is from carbonatite I and is characterized by low concentrations 510 of Sr, Fe, Mg, and Mn (Fig. 8). This variety contains the lowest contents of Ba, Pb 511 and ΣREE (6.55–241, 51.9–109, and 448–1221 ppm; on average 29.3, 70.8, and 876 512 ppm, respectively), but relatively high Y concentration of 92.5–281 ppm (mean = 206 513 514 ppm). The Zn, Th, As, and U abundances are low. The chondrite-normalized REE 5/15 patterns of Cal-I exhibit a gentle negative slope with lower (La/Yb)_N and (La/Nd)_N 516 ratios of 1.41–5.34 and 0.37–1.11, respectively (Fig. 10a). In addition, the δEu , δY 518 respectively) (Fig. 10b) (Bau, 1996; Chakhmouradian et al., 2016b).

519 Calcite-II (Cal-II) occurs in carbonatite I, but differs from Cal-I in the presence 520 of baryte and bastnäsite-(Ce) inclusions (Fig. 4h). Compared to Cal-I, Cal-II contains, elevated Sr, Mg, Fe, Mn concentrations (mean = 0.009, 0.005, 0.007, and 0.011 521 a.p.f.u.; Table S2 and Fig. 8) but distinctly higher REE contents (747-2144 ppm; 522 mean = 1359 ppm) and LREE-rich chondrite-normalized REE patterns with higher 523 $(La/Yb)_N$ and $(La/Nd)_N$ values (4.72-35.9 and 0.5-2.2, respectively; Fig. 10a). 524 525 Calcite-III (Cal-III) is found in carbonatite II and exhibits variable, but generally higher Sr concentrations (mean = 0.01 a.p.f.u.; Fig. 8) relative to Cal-I and Cal-II. 526 Y/Ho values of this variety (up to 34,7) are mostly within the CHARAC field and 527 528 only a few samples have higher, non-chondritic values (Fig. 10b). Cal-III is characterized by the highest REE contents (1011–2720 ppm; mean = 1740 ppm) with 529 the chondrite-normalized REE patterns varying over two orders of magnitude from La 530 to Lu ($(La/Yb)_N = 6.52-61.5$; avg. = 26.7) and $(La/Sm)_N$ ratios up to 8.30 (Fig. 10a). 531 532 In common with Cal-II, Cal-III hosts low concentrations of Y, Pb, Zn (mean = 119, 87.8, and 9.57 ppm, respectively), but slightly elevated Na content (mean = 265 ppm). 533 534 Calcite-IV (Cal-IV) samples from the contact zones are further divided into 535 calcites either from carbonatites or within carbonated nordmarkite and alkali granite. 536 Compared to other generations of calcite, Cal-IV has the lowest SO₃ and SrO contents, 537 but is more enriched in FeO, MgO, and especially MnO (Fig. 8). Moreover, this

538	variety has low levels of Ba, Pb and U (<155, <126, and <0.61 ppm, respectively), but
539	much higher REE contents (818–2775 ppm; mean = 1350 ppm).
540	Chondritic-normalized REE profiles of Cal-IV show LREE-rich patterns with respect
541	to chondrite ((La/Yb) _N ratios of 11.3), and chondritic Y/Ho and δ Eu values (27.6 and
542	0.87, respectively) (Fig. 10a).
543	The four generations of calcite exhibit continuous chemical variation, as shown
544	in composition scatter diagrams between Ca and other major elements (Fig. 9).
545	Strontium increases gradually from Cal-I to Cal-II and Cal-III (from 0.03 to 0.15
546	a.p.f.u.), but eventually drops in Cal-IV (0.941-0.978 a.p.f.u.) accompanied by more
547	variable Ca concentrations (Fig. 9a). Magnesium is inversely correlated with Ca ($R^2 =$
548	0.615) where Mg concentrations steadily increase from Cal-I to Cal-IV (Fig. 9b); the
549	same is true of Fe and Mn with respect to Ca ($R^2 = 0.641$ and 0.665, respectively; Fig.
550	9c and d).
551	

552 Phlogopite

Representative analyses of the chemical compositions and site assignments of 553 mica are listed in Table S4. Formula calculations for mica are based on an ideal 554 trioctahedral formula (XY₃[Z₄O₁₀][OH, F, Cl]₂) and the results are normalized to 7 555 556 (tetrahedral [Z] plus octahedral [Y]) due to possible vacancies on interlayer sites [X]. 557 All micas from the HZKS-1 drill are classified into phlogopite core $([K_{0.954}(Al_{0.203}Mg_{2.332}Fe_{0.368}Ti_{0.021}Mn_{0.007})(Al_{0.466}Si_{3.534}O_{10})(F,OH)_2]) \ \ based \ \ on \ \ high$ 558

Mg contents (1.991–2.750 a.p.f.u.) and low tetrahedral Fe³⁺ (<0.001 a.p.f.u.) (Foster, 559 560 1960). These phlogopites generally contain low Ba, Ca, and Na (<0.016, <0.009, and <0.033 a.p.f.u., respectively), and high F and K (mean = 1.369 and 0.963 a.p.f.u., 561 562 respectively), and low Na contents (0.009-0.033 a.p.f.u.). A strong negative correlation between Mg and Si towards higher Al indicates the effect of the 563 phlogopite–eastonite substitution (Mg²⁺ + Si⁴⁺ \leftrightarrow ^{IV}Al³⁺ + ^{VI}Al³⁺, Fig. 11a). In order 564 to investigate the degree of halogen enrichment, intercept values, IV(F), IV(Cl), and 565 IV(F/Cl), are calculated by equations: $IV(F) = 1.52X_{nbl} + 0.42X_{ann} + 0.20X_{sid} - 0.00X_{sid}$ 566 $\log(X_{F}/X_{OH})$; $IV(Cl) = -5.01 - 1.93X_{phl} - \log(X_{Cl}/X_{OH})$; and IV(F/Cl) = IV(F) - IV(Cl)567 (Munoz, 1984). Three populations of phlogopite can be identified from major element 568 compositions and halogen intercept values: (1) One population in carbonatite I (Phl-I), 569 570 (2) a main core-rim population from contact zones (Phl-II), and (3) accessory phases 571 in nordmarkite (Phl-III). Phl-I is discriminated by the lowest F contents (0.927–1.198 a.p.f.u.), with high 572 levels of Ti (up to 0.052 a.p.f.u., Fig. 11b and c), intermediate Mg# (88-95) and Na 573 (0.009–0.021 a.p.f.u.), and minor Mn (Fig. 11f). The intercept values of IV(F), IV(Cl) 574 575 and IV(F/CI) of this group range from 1.85 to 2.10, -2.60 to -3.90, and 4.59 to 5.86, 576 respectively. In comparison, Phl-II hosts higher Na (up to 0.033 a.p.f.u.) and F 577 (1.013–1.717 a.p.f.u.) (Fig. 11g-i), and the lowest Mn and Ti contents (Fig. 11b-d). Moreover, they are further distinct in having variable Mg# (85-97) with an 578 Mg-Si-rich core (Phl-IIC) and an Al-rich rim (Phl-IIR) (Fig. S5). Phl-III is 579

and 0.009 a.p.f.u., respectively), intermediate IV(F/Cl) (4.64–5.81), and the highest

583 Mg# (94–98) and Mn contents (Fig. 11b–f).

584

585 Amphibole

The results of analyses of amphibole are summarized in Table S5. Formula 586 calculations for amphibole are based on 24 oxygens and (O, OH, Cl) = 2 as 587 $[K_{0.282}Na_{0.464}][Na_{1.503}Ca_{0.445}][Mn_{0.037}Mg_{2.799}Fe^{2+}_{1.161}Fe^{3+}_{0.629}][Al_{0.124}Ti_{0.008}Si_{7.925}]O_{22}(F,$ 588 OH)2. In general, the chlorine content is minimal, but all amphiboles contain 589 significant F, Fe, Mg, Na (0.46-1.73, 1.02-2.06, 1.47-2.21, and 2.63-3.80 a.p.f.u., 590 respectively), and measurable K (0.09–0.37 a.p.f.u.) and Ca (0.13–1.91 a.p.f.u.) 591 contents. Most amphiboles have high Mg/(Mg + Fe²⁺) ratios (0.58–1.00) and 592 octahedral Na contents (Na_B >1.50 a.p.f.u.), and are therefore classified as 593 magnesioriebeckite-magnesioarfvedsonite (Leake et al., 1997). Classification 594 595 diagrams show that amphibole in carbonatite I is mainly magnesioriebeckite, whereas 596 those in carbonatite II and contact zones mainly belong to magnesioarfvedsonite (Fig. S6). 597

598

599 Clinopyroxene

SCRI

600	Major compositions of clinopyroxenes are listed in Table S6. The measured
601	compositions range from $Ae_{97}Di_3Hd_0$ to $Ae_{52}Di_{11}Hd_{37}$, where clinopyroxenes in
602	contact zones are aegirine with some aegirine-augite, and the main Ca-rich
603	aegirine-augite occurs in carbonatite II (Fig. S7). All samples have significant Mg (up
604	to 0.322 a.p.f.u.) with minor Mn, Al and Ti contents (up to 0.026, 0.053, and 0.039
605	a.p.f.u., respectively).
606	
607	C-O isotopic compositions of calcite in carbonatite
608	The C and O isotopic compositions of calcite separates from carbonatite I, II, and
609	the contact zones are presented in Table 2. The results show that $\delta^{13}C_{\text{VPDB}}$ values
610	range from –7.1 to –6.2 ‰ and $\delta^{18}O_{VSMOW}$ values are between 5.4 and 11.0‰, which
611	mostly plot in the box of primary igneous carbonate (PIC, Taylor et al., 1967) (Fig.
612	12). A number of data from Cal-IV in carbonatite II samples are located to the right of
613	the PIC field. The O isotope ratios show a progressive increasing trend from Cal-I to

614 Cal-III with a range from 6.6 to 9.2‰ but, in turn, decrease to 8.1‰ in Cal-IV.

615

616 In-situ Sr-C isotopic compositions recorded in various calcites

617 The measured results of in-situ Sr–C isotopic analyses of calcites are listed in 618 Tables 3 and 4. The low ⁸⁷Rb/⁸⁶Sr ratios (<0.00008) and high Sr contents (over 7000 619 ppm) of all varieties of calcite indicate that the measured ⁸⁷Sr/⁸⁶Sr ratios can 620 accurately record the initial Sr isotopic composition due to negligible effects from Rb 621 decay. The radiogenic Sr isotopic compositions obtained in the earlier-crystallizing Cal-I and Cal-II are similar to those of carbonatites (0.70605-0.70632, mean = 622 0.70613, Wang et al., 2001; Xu et al., 2003), as expressed by ⁸⁷Sr/⁸⁶Sr ratios of 623 0.70605 - 0.70623 (mean = 0.70613) and 0.70605 - 0.70636 (mean = 0.70617). 624 respectively (Fig. 13). In contrast, the Sr isotopic ratios of Cal-III and Cal-IV yield 625 high and variable ranges from 0.70602 to 0.70670 (mean = 0.70629) and 0.70605 to 626 0.70660 (mean = 0.70626), respectively, indicating more radiogenic Sr characteristics 627 628 (Fig. 13). In addition to Sr isotopic composition, the four generations of calcite are further 629 compositions (Table 4). A progressive distinguished by their in-situ C isotopic 630 decrease in $\delta^{13}C_{V-PDB}$ occurs from a high-value $\delta^{13}C$ of Cal-I (mean = -5.55‰, n = 26) 631 downward to relatively low δ^{13} C values (mean = -6.26‰, n = 24) in Cal-II, followed 632 by more depletion in δ^{13} C in Cal-III and Cal-IV (up to -9.97 and -8.61%), 633 respectively), which slightly deviates from mantle values (Fig. 17d). This feature 634 differs from the bulk C isotopic compositions analyzed for calcite (Niu et al., 2005; 635 Xu et al., 2002; Hou et al., 2006, 2015) and carbonatites in previous studies (Hou et 636 al., 2015). 637 638

639 **DISCUSSION**

640 Origins of calcite and silicate minerals in various carbonatites at Maoniuping

641 Most known carbonatites are derived from carbonate-silicate magma via liquid immiscibility or fractional crystallization (Wyllie, 1998) and have the ability to carry 642 643 REE, Sr and Ba (Veksler et al., 1998; Song et al., 2016). Carbonatitic magmas are 644 capable of rapid ascent owing to their low viscosity and density (Walter et al., 2021), and could intrude shallow crust as carbonate cumulates (Mitchell, 2005; Xu et al., 645 2010). Carbonatites from drill hole HZKS-1 at Maoniuping are carbonate cumulates 646 and are similar to those that formed by carbonatitic melts (Xu et al., 2010; 647 Chakhmouradian et al., 2016a) and are composed of large proportions of calcite and 648 minor silicate minerals, such as mica, clinopyroxene and amphibole (Reguir et al., 649 2009, 2012). Textural and chemical evidence suggest that carbonatite I and II derived 650 from different carbonatitic magma stages, and record the evolution of carbonatitic 651 magma. Reliable interpretation of the evolution of carbonatitic magma and 652 assessment of its relationship to REE enrichment and mineralization, requires 653 654 discrimination of the origins of the different generations of calcite and silicate minerals in various calcite carbonatite stages. 655

656 Calcite is the principal constituent and accounts for more than 90 vol.% of 657 carbonatite in some samples. Considering that calcite is easily amenable to textural 658 re-equilibration and subsolidus hydrothermal overprinting at low temperature and 659 pressure (Barker, 2001; Schultz et al., 2013; Chakhmouradian et al., 2016a), 660 determining the genesis of calcite requires an approach that combines petrographic observations, mineral chemical and isotopic compositions. The four generations of 661 calcites show near-CHARAC Y/Ho ratios and PIC-like C-O compositions. These 662 characteristics clearly indicate magmatic origin of all generations of calcite, and 663 significantly differ from those of hydrothermal or metasomatic calcites, either of 664 which has wide Y/Ho ranges (15–55) and variable $\delta^{13}C - \delta^{18}O$ values completely 665 outside the PIC field (Figs. 8a and 11b; Bau, 1996; Taylor, 1967). Among four 666 generations, Cal-I is considered as the early-crystalizing product from initial 667 carbonatitic melt, owing to: (1) the lowest Sr, Na, Ba, Pb and REEs with relatively 668 high Y contents, similar with those of early-crystalizing calcite reported in literature 669 (Table S3) (Hornig-Kjarsgaard, 1998; Bühn et al., 2001; Chakhmouradian et al., 670 2016b); (2) the gently sloping chondrite-normalized REE patterns with the lowest 671 $(La/Yb)_N$ and $(La/Nd)_N$ ratios (Fig. 11a); (3) the highest in-situ $\delta^{13}C$ values of -5.5%, 672 indicating a lack of modification by fluid-rock interaction, Rayleigh isotopic 673 fractionation, or decarbonation (Fig. 12); (4) the limited carbonatite-like ⁸⁷Sr/⁸⁶Sr 674 ratios; and (5) polygonal mosaic textures and near unity δCe values (0.98–1.04) and 675 slight Eu deficiency ($\delta Eu = 0.89-0.93$) of their host carbonatite I (Fig. 4a and c). 676 These features are evidenced by Cal-I, and, as the early melt is expected to have the 677 678 lowest La, REE contents and low (La/Yb)_N, the later calcite generations (Cal-II, Cal-III and Cal-IV) are inferred to have crystalized from more evolved carbonatite 679 680 magma.

681	Phlogopite is the essential K-bearing silicate mineral in calcite carbonatites and
682	is enriched in Mg, Ti but poor in Fe^{2+} (Table S3 and Fig. 11). Strong textural and
683	chemical evidence suggest that Phl-I is the primary magmatic phase that crystallized
684	directly from carbonatitic melt (Reguir et al., 2009; Massuyeau et al., 2015). In
685	contrast, Phl-II in contact zones cannot be ascribed to magmatic genesis owing to
686	their high abundance and distribution alongside aegirine/aegirine-augite and
687	microcline (Fig. 6a). These features reflect an increased Si activity in the carbonatite
688	magma, which is opposite to the general trend of decreasing SiO ₂ contents and silica
689	activity during fractionation of carbonatitic magma (Barker, 2001; Krasnova et al.,
690	2004; Lee et al., 2004; Reguir et al., 2012; Massuyeau et al., 2015; Weidendorfer et al.,
691	2017; Giebel et al., 2019a, b and references therein). Thus, we infer that Phl-II is a
692	metasomatic product due to interaction between carbonatite melt and silicate rocks
693	(nordmarkite).
694	Amphibole (riebeckite and arfvedsonite) and clinopyroxene (aegirine-augite and
695	aegirine) are important ferromagnesian silicates in carbonatites in the HZKS-1 drill
696	core, and are chemically dominated by sodic and sodic-calcic varieties. As riebeckite
697	occurs in carbonatite I and shows euhedral crystals enclosed in Phl-I, they are

698

interpreted as primary silicates that crystallized from early carbonatitic melts (Reguir 699 et al., 2012; Giebel et al., 2019a, b). Arfvedsonite, aegirine and aegirine-augite occur both inside carbonatite II and at contact zones (Figs. 5 and 6) as euhedral-subhedral 700 701 crystals and are compositionally distinct from those found in nordmarkite (Figs. S5

- nordmarkite. Thus, we suggest that these minerals formed similar to Phl-II (see above)
- 704 by the interaction between carbonatitic melt and nordmarkite.
- 705 The Cal-I-, Phl-I- and riebeckite-bearing carbonatite I exhibits equiangular 706 texture with the lowest REE contents, LREE/HREE ratios and near-chondritic Y/Ho values (Figs. 7 and 10). Both petrographic features and chemical compositions of 707 708 carbonatite I samples are consistent with those in primary carbonatites that are confirmed to derive from early carbonatite melt, such as Kalkfeld, Namibia, Homa 709 Moutain, Kenya and Lackner Lake, Canada (Bühn et al., 2001; Chakhmouradian et al., 710 2016a). Therefore, the carbonatite I from the deep drill core (Fig. 3g) could be the 711 best representation of initial carbonatites at Maoniuping. Such carbonatites should 712 713 inherit their compositional signatures from early magma, showing enrichment of alkalis, fluorine and REEs (>3000 ppm) but relatively poor in SiO₂ (<5 wt.%), MgO 714 (<0.5 wt.%), FeO (<1.0 wt.%) and especially in P₂O₅ (<0.5 wt.%). 715
- 716

717 Fractional crystallization of carbonatitic magma leading to REE enrichment

As discussed above, the four calcite generations are inferred to be of magmatic origin and carbonatites lack hydrothermal modification. Hence, it is feasible to trace magma evolution using the chemical characteristics of calcite, fluorite, phlogopite, and amphibole combined with the bulk chemistry of carbonatites. Although the major and trace element compositions of investigated samples cannot be used to diagnose

723	the composition of the parental silicate-carbonate magma, they can be invoked to test
724	carbonatitic evolution and degree of magma fractionation (Horning-Kjarsgaard, 1998;
725	Bühn et al., 2001; Chen & Simonetti, 2013; Chakhmouradian et al., 2016b; Su et al.,
726	2019; Anenburg et al., 2020b, 2022; Gao et al., 2021). To interpret their degree of
727	evolution, equilibrium and fractional (Rayleigh) crystallization were modelled and
728	represented in terms of La and Ce concentrations owing to their overwhelming
729	abundances compared to the other REEs (Pr-Lu). In our model calculations, partition
730	coefficients for La and Ce between calcite/fluorite and carbonatitic melt $(D_{calcite/melt}^{REE})$
731	and $D_{fluorite/melt}^{REE}$) were obtained from Chebotarev et al. (2019, 2021), determined by
732	experiments at 650-900 °C, whereas those for amphibole and carbonatitic melt
733	$(D_{amphibole/melt}^{REE})$ are from Bottazzi et al. (1999) and Reguir et al. (2012). Fractional
734	crystallization increments from $F = 1.0$ to 0.1 were employed with initial whole rock
735	compositions based on carbonatite I and appropriate calcite:amphibole:fluorite ratios
736	ranging from 0.9:0.1:0 to 0.85:0.05:0.1 based on observed mineral proportion from
737	carbonatite I and II. Results of the modal calculation for equilibrium and fractional
738	crystallization are illustrated in Fig. 14. In general, both differentiation models can
739	effectively facilitate enrichment of REE due to $D_{calcite/melt}^{REE}$ values less than unity
740	(Chebotarev et al., 2019, 2021). However, calcite crystallized from equilibrium
741	crystallization would have a slow increase in La and Ce — that is, enrichment in
742	equivalent REE contents through equilibrium crystallization requires higher degrees
743	of magma fractionation compared to Rayleigh crystallization (Fig. 14a and b). In
744 addition, the equilibrium model does not reconcile adequately with the measured La and Ce concentrations, as a significantly higher degree of magma fractionation (F =745 746 ~ 0.1) could not produce compositions extending into the higher ranges of La and Ce 747 concentrations (up to 300 and 1200 ppm, respectively) (Fig. 14a). Modifying the proportions of calcite, amphibole and fluorite would yield similar results (F >0.85 for 748 enrichment in La and Ce in Cal-III), indicating crystallization of amphibole and 749 fluorite have no prominent effect on LREE enrichment during carbonatitic magma 750 differentiation. In contrast, the calculated results for the Rayleigh fractional 751 crystallization model not only overlap the measured data well, but display a 752 near-linear positive correlation in which Ce rapidly enriches with La contents with the 753 degree of fractional crystallization ranging from F = 0.95 to 0.4 (Fig. 14c and d). Of 754 755 note, the choice of modal ratio (calcite:amphibole:fluorite = 0.9:0.1:0 to 0.85:0.05:0.1) or increasing fractional crystallization increments ($\Delta F = 10\%$, 5%, and 1%) cannot 756 change the modal tendency between La and Ce contents, except increasing the 757 enrichment rate of REE concentrations in residual melts (Fig. 14b-d). For instance, 758 enrichment to 1200 ppm Ce and 600 ppm La in crystallizing Cal-III requires a 759 760 significant degree of fractionation (F = ~ 0.8 , $\Delta F = 1\%$, Fig. 16b), which is much greater than that modelled for intense fractionation with F = 0.5 to 0.6 at $\Delta F = 5\%$ to 761 762 10% (Fig. 14c and d).

In general, REE-rich carbonatites are seemingly derived from more evolved
melts (Chamouradian & Zaitzev, 2012; Walter et al., 2021; Anenburg et al., 2022) and

765 this has been suggested not only by experiments (Anenburg et al., 2020b; Chebotarev et al., 2019, 2022), but in natural cases involving several African carbonatite 766 767 complexes (Bühn et al., 2001), Bayan Obo and Huayangchuan, China (Yang et al., 2019; Gao et al., 2021), Kaiserstuhl, German (Walter et al., 2018, 2020; Giebel et al., 768 2019b) and elsewhere. Our modal calculations suggest that with equilibrium 769 crystallization it is much more difficult to produce the extremely high La and Ce 770 contents in Cal-III at Maoniuping. Fractional crystallization, instead, is not only more 771 suitable for explaining the measured trace element data, but also consistent with the 772 textural evolution of calcites. Early REE-bearing phases (i.e., Cal-I, La = 85 ppm, Ce 773 = 240 ppm, F = 0.95) crystallized at deep levels would drive the residual melts toward 774 a composition saturated with REE-bearing phases (Cal-II, La = 190 ppm, Ce = 580 775 776 ppm, F = -0.8), and consequently evolve to REE-rich and super coarse-grained products (Cal-III, La = 340 ppm, Ce = 830 ppm; F = 0.5 to 0.6). Thus, it is plausible 777 that carbonatite II derived from rather residual carbonatite melt, causing its pegmatoid 778 779 texture. Systematic and continuous variations in major compositions (Fig. 9) and 780 positive correlation between La and Ce, (La/Yb)_N and REE, from Cal-I to Cal-IV (Fig. 781 15), clearly demonstrate that both chemical evolution and REE enrichment can be 782 attributed to protracted fractional crystallization in an individual melt pulse. This is in 783 part because neither mixing by carbonatitic melt batches with counterpart silicate 784 melts nor multiple carbonatitic melt pulses could provide progressive variations in investigated REE data, resulting in scattered, extended and even uncertainly 785

link-absent trends, as reported by experimental results (Foley et al., 2009) and natural
cases in the Oka (Chen & Simonetti, 2013) and Aley carbonatites (Chakhmouradian
et al., 2015).

789 Although intense fractional crystallization is able to generate the observed REE enrichment in calcites and host carbonatites, our modal calculations indicate that 790 progressive calcite accumulation from carbonatitic melt can only yield constant 791 (La/Yb)_N ratios between 15 and 18, conflicting significantly with steepening of REE 792 profiles and high (La/Yb)_N values from 3.09 to 26.8 (Fig. 10a). Given low 793 $D_{calcite/melt}^{HREE}$ values (Ho-Lu from 0.05 to 0.07, Chebotarev et al., 2019) and 794 abundant proportions of amphibole, phlogopite and elinopyroxene in carbonatite II, 795 we suggest that increased (La/Yb)_N ratios observed from Cal-I to Cal-III resulted from 796 silicate crystallization. Owing to their high distribution coefficients for HREE 797 798 (Chebotarev et al., 2021), their crystallization in early carbonatite melt could 799 sequestrate HREE and lead to an increase of (La/Yb)_N in residual melt and derived calcites (e.g., Cal-II and Cal-III). As a result, the positive correlation between 800 801 (La/Yb)_N and REE concentrations is controlled by the relative proportions of calcite 802 and silicates during fractional crystallization of carbonatitic magma (Fig. 15b). Such 803 increased correlation between LREE-HREE fractionation and REE contents probably 804 occurs in P-poor carbonatitic system like Maoniuping, because early-crystalizing 805 apatite/monazite in P-rich systems leads to REE depletion during fractional 806 crystallization, resulting in REE-poor later-stage carbonatites (i.e., Magnet Cove,

807	USA, Italite Lake, Callada, and Kerlinasi, Talizania, Milaoya, Chilla, Hg. 15,
808	Chakhmouradian et al., 2016b). Additionally, the discernable relationship between
809	redox-sensitive element pairs, including a decrease in δCe with constant unity of δEu
810	from Cal-I to Cal-III, clearly supports an increase in O ₂ fugacity (fO ₂) during magma
811	fractionation (Fig. 15a and d), similar to global carbonatites, such as Palabora
812	(Horning-Kjarsgaard, 1988; Giebel et al., 2017) and Aley (Chakhmouradian et al.,
813	2015).
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815 Sr-C isotope decoupling and REE mineralization during contamination by
816 silicate wall rocks

The Maoniuping carbonatites have been interpreted to originate from fertile 817 sub-continental lithospheric mantle (SCLM, Hou et al., 2015), according to 818 consistently high whole-rock 87 Sr/ 86 Sr ratios (>0.7055) and mantle-like δ^{13} C- δ^{18} O 819 values (-6.6% to -7.0% and 6.7% to 8.4%) (Wang et al., 2001; Xu et al., 2003; Hou 820 et al., 2006; Liu & Hou, 2017). However, our in-situ isotopic measurements for four 821 generations of calcite from carbonatite I and carbonatite II provide different 822 constraints on evolution, not only from radiogenic Sr, but from stable C isotopic 823/ 824 composition. The in-situ Sr-C isotopic variations indicate that the earlier-crystalizing Cal-I and Cal-II have low 87 Sr/ 86 Sr ratios (mean = 0.70613 and 0.70617) but relatively 825 high δ^{13} C values (mean = -5.55 ‰ and -6.26 ‰), whereas Cal-III and Cal-IV are 826 enriched in ⁸⁷Sr/⁸⁶Sr ratios (up to 0.70670 and 0.70660) and have a greater depletion 827

in δ^{13} C (drop to -9.97 ‰ and -8.61 ‰) (Tables. 3 and 4). Theoretically, the variable 828 Sr isotopic composition in Cal-III and Cal-IV could be caused by either hydrothermal 829 reworking, external fluid addition, or contamination with wall-rock. However, 830 831 hydrothermal reworking can be ruled out as there is little compositional and petrological evidence for fluid overprinting both in our observations and previous 832 studies at Maoniuping (see above and Xie et al., 2015, Liu & Hou, 2017; Zheng & Liu, 833 2019). External fluid addition (e.g., by meteoric water or subduction-related fluids) 834 could lead to Sr isotopic re-equilibration (Ying et al., 2020; Wei et al., 2022), but it is 835 also expected to cause textural re-equilibration of calcite grains and wide chemical 836 variation, especially in Y/Ho ratios (18.4–46.2), extreme depletion in Eu (δ Eu drops 837 to 0.3) and low Sr contents (Chakhmouradian, 2016b). However, none of these 838 features are observed in any of the studied calcite generations and hence, external 839 fluid addition is unlikely at Maoniuping. In contrast, contamination with silicate 840 wall-rock (nordmarkite) during carbonatitic magma ascent is most likely the essential 841 factor in variable Sr isotopic composition (Fig. 13). Under this scenario, the high 842 ⁸⁷Sr/⁸⁶Sr ratios in Cal-III and Cal-IV require a more radiogenic Sr source. Considering 843 844 that both nordmarkite counterparts and wall-rock feldspars are capable of 845 incorporating significant levels of radiogenic Sr (87 Sr/ 86 Sr ratios = 0.70638–0.70686 846 and 0.70704, respectively, Xu et al., 2003; Hou et al., 2006; Liu & Hou, 2017) and hence, they are likely contributors to the variable ⁸⁷Sr/⁸⁶Sr ratios. This is also 847 848 supported by petrographic evidence of metasomatic microcline in carbonatite II at the

849	contact zones (Fig. 6a), and is similar to the situation reported in some well-studied
850	carbonatites, such as Caotan in China (Wei et al., 2020), Nolans Bore in Australia
851	(Anenburg et al., 2020a), Blue River in Canada (Mitchell et al., 2017), Vuoriyarvi in
852	Russia (Fomina & Kozlov, 2021), Kaiserstuhl in Germany (Walter et al., 2018; Giebel
853	et al., 2019b), Kieshöhe and Gross Brukkaros in Namibia (Walter et al., 2022, 2023)
854	Contamination by nordmarkite cannot be identified easily by bulk Sr-Nd-Pb
855	isotopic compositions because their much higher elemental concentrations in
856	carbonatites would buffer against change in the isotope systems during contamination
857	with silicate rocks (Bell & Tilton, 2002). However, our in-situ Sr-C isotopic data can
858	be used to track the processes of contamination by more radiogenic Sr isotopic
859	composition in Cal-III and Cal-IV relative to early-crystalizing Cal-I and Cal-II. To
860	explain the level of contamination responsible for the mixed Sr isotopic compositions,
861	a Monte Carlo simulation has been conducted by assuming that initial ⁸⁷ Sr/ ⁸⁶ Sr ratios
862	are 0.70605 for carbonatite according to our analyses of Cal-I and 0.70686-0.70700
863	for nordmarkite (Xu et al., 2003; Hou et al., 2006). Based on whole-rock results of our
864	and previous studies, Sr contents are selected from 8000 to 25000 ppm for
865	carbonatites (Hou et al., 2015) and 500 to 3000 ppm for nordmarkite (Wang et al.,
866	2001), with initial REE contents of 1000 to 3000 ppm (Wang et al., 2001; Hou et al.,
867	2006) and 200 to 800 ppm (Xu et al., 2003; Hou et al., 2015) for carbonatite and
868	nordmarkite, respectively. A total of 4×10^8 simulations were performed by randomly
869	selecting input values from the defined parameters. In the modal calculation (Fig. 16),

a schematic trend in ⁸⁷Sr/⁸⁶Sr ratios relative to Sr and REE contents (Fig. 17a) 870 illustrates that as much as 40% (± 10%) of the Sr isotopic budget of Cal-III- and 871 Cal-IV-bearing carbonatite was contributed by wall-rock nordmarkite. 872 Coupled with increased ⁸⁷Sr/⁸⁶Sr ratios, in-situ C isotopic compositions display 873 discriminable variation in which initial high δ^{13} C values of early-crystalizing Cal-1 874 5.5‰) largely dropped to low δ^{13} C values in Cal-III and Cal-IV (up to -9.9% and -875 8.7%, respectively, Fig. 17d). Theoretically, the excursion in in-situ C isotopic 876 composition could be induced by fractional crystallization of magma, addition of 877 sedimentary carbonate or meteoric fluids, or interaction between carbonatitic magma 878 and wall rocks (Deines, 1992; Santos & Clayton, 1995; Ray et al., 1999). Since 879 fractional crystallization of magma, addition of sedimentary carbonate or meteoric 880 fluids would result in a positive shift in δ^{13} C (Santos & Clayton, 1995; Chazot et al., 881 2003; Ray & Ramesh, 2000), these three hypotheses do not reconcile the lowering of 882 δ^{13} C values measured in Cal-III and Cal-IV (Fig. 17b). By contrast, contamination via 883 interaction between carbonatitic melt and wall rocks is probably the reason for C 884 isotope deficiency in Cal-III and Cal-IV, because C isotopic composition would 885 886 change as a result of decarbonation reactions and re-equilibration between 887 carbonatitic melt and CO_2 by the following reactions: 888 Carbonatite melt + Quartz (*nordmarkite*) \rightarrow Aegirine + Calcite + CO₂ (*fluid/gas*) 889 $CaFe(CO_3)_2 + 0.5Na_2CO_3 + 0.25O_2 + 2SiO_2 = NaFeSi_2O_6 + CaCO_3 + 1.5CO_2$ 890 (reaction 1)

+ $8SiO_2 =$ (reaction 2) site + CO_2

898 $3Ca(Mg,Fe)(CO_3)_2 + H_2O + KAlSi_3O_8 = K(Mg, Fe)_3AlSi_3O_{10}(OH)_2 + 3CaCO_3 + CaCO_3 + CaCO$

Carbonatite melt + K-feldspar (nordmarkite) \rightarrow Phlogopite + Calcite

899 3CO₂

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 $(gas/fluid) + O_2(gas)$

 $5CaFe(CO_3)_2$

(fluid/gas)

(reaction 3)

Contamination-induced interaction between carbonatite melts and silicate rocks 901 902 could trigger reactions 1-3. Similar modification of silicate rocks by carbonatitic melts has been referred as "antiskarn" metasomatism (Anenburg & Mavrogenes, 2018) 903 and reflects mutual interaction during contamination. The Na₂CO₃ and H₂O 904 components in carbonatite melt at Maoniuping are supported by sodium-carbonate 905 906 and water-bearing melt and melt-fluid inclusions in calcite and fluorite (Xie et al., 907 2009; Zheng et al., 2019). Both soluble components, together with $CaFe(CO_3)_2$ and 908 Ca(Mg,Fe)(CO₃)₂ in carbonatite melt, react with SiO₂ and KAlSi₃O₈ of silicate wall 909 rocks to control the formation of aegirine and aegirine-augite, magnesioarfvedsonite, 910 phlogopite, and calcite. The chemical classification of arfvedsonite, aegirine, and 911 aegirine-augite show that they are similar to those of other carbonatites worldwide

891 Carbonatite melt + Quartz (*nordmarkite*) \rightarrow Magnesioarfvedsonite + Calcite + CO₂

 $1.5Na_2CO_3$

H₂O

 $4CaMg(CO_3)_2$

Na(Na,Ca)₂(Mg,Fe)₄FeSi₈O₂₂(OH)₂ + 7CaCO₃+ 12.5CO₂ + 2.75O₂

912 (Smith, 2007; Reguir et al., 2012). Their compositional evolution exhibits progressively enrichment in Ca and Mg from carbonatite I to contact zone and 913 914 carbonatite II (Figs. S6 and S7), indicating that prolonged contamination by silicate 915 wall rocks lead to removal of Ca and Mg from the carbonatite system by formation of magnesioarfvedsonite, aegirine-augite, and calcite (reactions 1-2). Moreover, 916 carbonatite melt would obtain Si and Al from silicate wall rocks and further expel 917 soluble components of Na, K and F (reaction 3) as evident not only by higher Na and 918 IV (F) values (Fig. 11g and i) but K-rich core and Al-rich rim (Fig. S5) in Phl-II. 919 920 Previous experiments showed that alkalis and fluorine are crucial ligands that can bond to REE as alkali-REE and fluorine-REE complexes and allow long-distances 921 mobilization in carbonatite systems (Song et al., 2016; Chebotarev et al., 2019; 922 923 Anenburg et al., 2020b); and the soluble alkali carbonate (Na₂CO₃) in carbonatite melt can significantly promote REE solubility under magmatic-hydrothermal conditions 924 (Anenburg et al., 2020b). Thus, it is plausible that both Na₂CO₃ consumption and 925 formation of abundant aegirine, aegirine-augite, arfvedsonite and phlogopite by 926 reactions 1-3 at Maoniuping serve to scavenge Na, K, and F from carbonatite melt, 927 928 which could destabilize alkali-REE and fluorine-REE complexes and thus probably 929 mhibit REE mobility and lead to accumulation not only in produced calcite (Cal-III 930 and Cal-IV) coupled with δ^{13} C depletion (Fig. 17c), but also as bastnäsite-(Ce) 931 intergrown with fluorite and silicate minerals in carbonatite II and the contact zone 932 (Figs. 5 and 6).

933	In addition, reactions 1–3 would contemporaneously cause strong
934	decarbonation, which would preferentially remove ¹³ C from carbonatitic melt by
935	release of CO ₂ (Chacko et al., 1991) and contribute to low $\delta^{13}C$ values in residual
936	carbonatitic melts and their derived calcites (Fig. 17d). Therefore, the gradual
937	depletion δ^{13} C values from Cal-I to Cal-III and Cal-IV, combined with increases of
938	⁸⁷ Sr/ ⁸⁶ Sr (Fig. 17b), robustly demonstrate protracted contamination by interaction
939	with wall rocks during carbonatite melt ascent. The processes of decarbonation are
940	delineated by experiments (Chacko et al., 2001; Anenburg & Mavrogenes, 2018) and
941	can be quantified using the equation for Raleigh volatilization (Valley, 1986): $\delta_f = \delta_i + \delta_i$
942	$1000 \times (F^{(\alpha-1)} - 1)$, where F is the C fraction remaining in the carbonatitic melt, and δ_f
943	and δ_i are the final and initial $\delta^{13}C$ values. The fractionation factor α between calcite
944	and CO ₂ was calculated by the experimental equation of Scheele and Hoefs (1992):
945	$10^{3} \ln \alpha_{\text{calcite-CO2}} = -3.46 \times 10^{6}/\text{T}^{2} + 9.58 \times 10^{3}/\text{T} - 2.72$. Assuming that δ^{13} C values of –
946	5.5% (Cal-I) and 7.5% (Cal-III) represent the initial and final C isotopic
947	compositions, respectively, and reaction temperatures are estimated at 500-800 °C
948	based on melt and fluid inclusion microthermometry, the fraction of carbon remaining
949	in reaction zones (F) is calculated to be ~0.67 (Fig. 17f), indicating that as much as
950	\sim 35% of initial CO ₂ could be degassed in contact zones due to carbonatite
951	melt-nordmakite interaction. As the degassed CO2 would not readily disperse but
952	could be retained in the magma (Harmer & Gittins; 1997; Russell et al., 2012), Walter
953	et al. (2021) proposed the concept that CO_2 release from decompressive immiscibility

954 of carbonatitic fluids could promote fracturing and drive carbonatite magma ascent. Equally, decarbonation of CO₂ by carbonatite melt-wall rock interaction and their 955 956 retention in carbonatite melt would act as a positive mechanism of increasing 957 buoyancy and drive magma rapid ascent. Contamination-induced CO₂ degassing, would also lead to volumetric expansion and fracturing through the generation of 958 over-pressures (Weidendorfer et al., 2019; Water et al., 2020, 2021). As a result, 959 continual contamination during carbonatite magma ascent is inferred to trigger 960 volumetric expansion, leading to fracturing and conduit formation for super 961 962 coarse-grained carbonatite II and simultaneously result in strain-induced, dynamic decomposition of carbonatite II (e.g., miarolitic and porous textures with undulatory 963 extinction and elongated calcite crystals Figs. 3k and 5i) and brecciation events both 964 965 at Maoniuping (see Fig. 6 in Liu et al., 2019a) and in most carbonatites worldwide (Smith et al., 2007; Elliott et al., 2018). 966

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968 A general model for REE enrichment and mineralization in carbonatites

969 Most known carbonatite-related REE deposits are closely associated with 970 metasomatized mantle carbonatite sources, with influence on REE mineralization 971 from magmatic evolution, and hydrothermal activity (Chamouradian & Zaitzev, 2012; 972 Smith et al., 2016; Anenburg et al., 2022). Numerous studies have contended that 973 metasomatized carbonatite sources influenced by subducted slab and including 974 components of recycled sediments contribute to the generation of initial REE-fertile

975	carbonatites (Hou et al., 2015; Xu et al., 2017; Çimen et al., 2018; Yaxely et al., 2022).
976	Subsequent late-stage hydrothermal reworking on such fertile carbonatites might be
977	responsible for REE mineralization in a number of deposits, including the Miaoya,
978	Cummins Range, and Bayan Obo carbonatites (Kynický et al. 2012; Smith et al., 2015;
979	Downes et al., 2016; Ying et al., 2020; Anenburg et al., 2022). Nevertheless, in a
980	global context, models for carbonatitic magma evolution and its impact on REE
981	mineralization are rare (Smith et al., 2016). REE accumulation caused by fractional
982	crystallization (Bayan Obo, China, Yang et al., 2019) and by contamination linked to
983	occurrence of mica and apatite (Kaiserstuhl, Germany, Giebel et al., 2019b) have been
984	inferred. However, the combination of different processes and their functions with
985	regards to REE enrichment and mineralization for the formation of carbonatite-related
986	REE deposits has not been examined in an integrated model.
987	In the case of Maoniuping, it is believed that mantle source became enriched in
988	REE by recycling of Neoproterozoic subducted marine sediments (Hou et al., 2015).
989	The Maoniuping carbonatites were likely derived from the fertile mantle source in the
990	Cenozoic (Hou et al., 2015; Liu et al., 2019a; Weng et al., 2022b) and experienced
991	further REE accumulation by magma evolution. Here, we reconcile petrographic and
992	geochemical data to propose a conceptual model for the REE enrichment and
993	mineralization during carbonatite magma evolution involving fractional

994 crystallization and contamination by silicate rocks (Fig. 18).

Initially a fertile carbonatite melt was derived from carbonate-silicate magma by liquid immiscibility which probably occurred at shallow-level crustal level as evidenced by low MgO contents, low (Mg+Fe)/Ca and petrographic evidence of Cal-I and carbonatite I. Extremely low viscosities of carbonatitic melts enable them to ascend rapidly (Jones et al., 2013; Kono et al., 2014). Owing to the lower density of carbonatitic melt (<2400 kg/m³, Jones et al., 2013; Massuyeau et al., 2023) than syenitic melts (~2500 kg/m³, Semenov & Polyansky, 2017), the initial carbonatite melt may be suspended in or surrounded by silicate melts, which could prevent carbonatitic melt from reacting immediately with basement rocks (Fig. 18a). When reaching liquidus temperatures (e.g., >900°C, Panina, 2005; Guzmics et al., 2011), the first-order carbonate phases (e.g., Cal-I) began to crystallize, with precipitation of minor, primary phlogopite (Phl-I) and riebeckite at a low degree of fractional crystallization (F = $\sim 0.95-0.90$) of carbonatite melt. These phases are expected to inherit the compositional characteristics of the initial carbonatitic melt, as indicated by the lowest REE contents (~800 ppm), low (La/Yb) ratios (~3.1), and primary Sr-C isotopic composition (87 Sr/ 86 Sr = 0.70613 and δ^{13} C = -5.5‰) (Fig. 18b1). Considering the high incompatibility of alkalis, REE and halogens (Chebotarev et al., 2019), calcite accumulation drove increased REE concentrations in residual melts and crystallization of alkali-bearing and REE-rich Cal-II and slightly enhanced formation of halogen-bearing phlogopite (Phl-I). The gradually increasing alkali and volatile

1015 contents and decompression, in turn, lowered the viscosities of carbonatitic melt

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1017 et al., 2023) as recorded by the low-strain texture in Cal-II (Fig. 18b2).

1018 While ascending to higher crustal levels, further fractional crystallization (F =0.7-0.8) of carbonatitic melt derived carbonate cumulates lead to enrichment of 1019 1020 alkaline, volatile, and REE components in residual melts. At the same time, contamination with silicate rock via metasomatic "antiskarn" interactions between hot 1021 1022 carbonatitic melt (>600 °C, Xie et al., 2009) and silicate rocks occurred. Such interactions caused incorporation of more radiogenic Sr from host rocks, and lowered 1023 1024 C isotopic ratios through decarbonation reactions. This triggered the formation of and 1025 aegirine and aegirine-augite, magnesioarfvedsonite inside phlogopite, carbonatites by a series of reactions between dissolved carbonates of Na₂CO₃, K₂CO₃, 1026 1027 and $Ca(Mg,Fe)(CO_3)_2$ and SiO_2 and Al_2O_3 from silicate rocks (Fig. 18c). The CO₂ generated by decarbonation reactions would not release instantaneously but be 1028 retained in carbonatitic melt, causing the development of overpressures, which drove 1029 a jet-like ascent (Fig. 18a; Walter et al., 2021). More importantly, wallrock 1030 contamination resulted in REE accumulation and mineralization, as evident by a 1031 1032 significant proportion of bastnäsite-(Ce) with calcite (Cal-III and Cal-IV) that contain 1033 highest REE concentrations (Fig. 6). Magmatic and magmatic-hydrothermal transport 1034 of the REE has been demonstrated to be as complexes with a wide range of ligands including F⁻, Cl⁻, CO₃²⁻, PO₄³⁻ and SO₄²⁻ (Migdisov et al., 2016), with F⁻ and PO₄³⁻ 1035 1036 specifically acting as depositional ligands in hydrothermal systems (Williams-Jones et

1037	al., 2012). Formation of fluorine-bearing silicates and fluorite could destabilize
1038	halogenated REE complexes in the melt, resulting in REE precipitation as
1039	bastnäsite-(Ce). Similarly, Anenburg et al. (2020b) demonstrated that alkalis were
1040	critical in magmatic-hydrothermal REE transport, although specific complexes have
1041	not yet been identified. Si and Al contributions from wall rocks and consumption of K
1042	and Na by contamination-triggered formation of alkali silicates may have resulted in
1043	reduced REE solubility, and then formation of bastnäsite-(Ce) (Fig. 18b3).
1044	As carbonatitic melt ascended, pronounced factional crystallization of
1045	carbonatite melt (F = $0.5-0.6$) lead to REE enrichment to significant levels,
1046	accompanied by calcite cumulates with pegmatoid texture and super large crystals
1047	(carbonatite II). Continuous contamination admixed radiogenic Sr and lowered $\delta^{13}C$
1048	isotope values ($\delta^{13}C = -7.5\%$ and ${}^{87}Sr/{}^{86}Sr$ ratios up to 0.70660) by further interaction
1049	and decarbonation CO_2 . These CO_2 releases leads to vigorous volumetric expansion
1050	that triggers the ubiquitous brecciation observed at Maoniuping, and likely affected
1051	calcites so that they exhibit stain-induced deformation, such as bent and segmented
1052	thick twinning lamellae in Cal-III (Fig. 18b4). Moreover, repeated interaction between
1053	REE-rich carbonatite melts and silicate rocks scavenge more alkalis and volatiles, and
1054	triggered formation of a large proportion of the bastnäsite-(Ce), fluorites, baryte, and
1055	alkaline silicates (Fig. 18b4). The case studies cited above indicate that this may be a
1056	widespread mechanism in P-poor carbonatites.

1057

1058 CONCLUSIONS

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1059 The Maoniuping carbonatite-related REE deposit in the southwestern China is 1060 characterized by medium- to coarse-grained carbonatites in the lower part of drill core, which transition gradually upward into super coarser-grained carbonatite with 1061 1062 pegmatoidal texture. Four well-defined generations of calcite and compositional variation of phlogopite, amphibole and clinopyroxene in two types of carbonatites 1063 1064 show a systematic increase of REE contents, but progressively decoupled Sr-C isotopic compositions. Based on petrographic observations and geochemical data, we 1065 1066 interpret the pronounced REE enrichment in Maoniuping carbonatites as a result of 1067 protracted fractional crystallization of carbonatite magma, whereas more radiogenic Sr and gradually depleted C isotopic compositions are attributed to interaction with 1068 1069 silicate rocks and decarbonation of CO_2 loss. These new findings support an 1070 which REE mineralization could be integrated model in triggered by 1071 contamination-induced interaction between carbonatitic melt and silicate rocks. From this study, we propose that magmatic processes may play a more important role in 1072 REE enrichment and mineralization than previously recognized in carbonatites 1073 1074 worldwide, which could be a significantly factor in the formation of giant carbonatite-related REE deposits. 1075

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1594 TABLE CAPTIONS

1595 Table 1. Whole-rock chemistry of carbonatite I, II, nordmarkite, and alkali granite

		Carbo	antit											Carbo	natite II		
	Ro	e I															
	ck																
	Sa	нгк	нzк	нzк	нzк	нzк	нzк	нzк	нzк	HZKS	нzк	HZKS	нzк	нzк	нтк	HZKS	
	mp	S-1-	S-1-	S-1-	S-1-	S-1-	S-1-	S-1-	S-1-	-1-10	S-1-	-1-10	S-1-	S-1-	S-1-	-1-18	
	le	3	6	7	17	22	2B	5A	5B	Α	2A	В	31	4	13	Α	
	No)		
	•											/					
	wt.												>				
	%										1		7				
	SiO	0.8	0.4	0.8	0.29	0.35	3.09	0.45	0.44	0.78	5.00	####	###	2.3	5.45	9.32	
	2	3	0	2								#	#	2			
	TiO	0.0	0.0	0.0	0.00	0.02	0.00	0.01	0.01	0.04	0.10	0.08	0.08	0.0	0.02	0.08	
	2	0	1	0						\succ	r			9			
	AI_2	0.0	0.2	0.0	0.01	0.04	0.19	0.06	0.06	0.49	1.98	3.38	3.38	0.3	0.07	0.37	
	O ₃	1	2	0					Y					7			
	TF	0.6	0.7	0.6	0.64	0.87	0.50	0.92	0.92	0.61	0.95	0.61	0.79	1.5	0.56	1.52	
	e ₂	4	7	5										3			
	O ₃					$\langle \rangle$	\bigvee										
	Fe	0.4	0.4	0.4	0.43	0.58	0.39	0.69	0.66	0.31	0.50	0.44	0.38	0.3	0.22	0.42	
	0	3	9	4		Y								2			
	Mn	0.9	0.9	0.9	0.97	0.93	0.55	0.83	0.82	0.44	0.60	0.67	0.81	0.4	0.41	0.42	
	0	8	4	8										2			
	Mg	0.2	0.6	0.2	0.26	0.65	0.14	0.20	0.20	0.13	0.17	0.18	0.29	0.1	0.21	0.14	
	0	5	9)	4										3			
	Ca	###	, ###	###	###	###	###	###	###	####	###	####	###	###	###	####	
	ø	##	##	##	##	##	##	#	##	#	##	#	#	##	##	#	
$\boldsymbol{\checkmark}$	Sr	1.3	1.5	1.3	1.27	1.78	2.10	1.19	1.16	0.88	1.16	0.96	1.38	1.5	1.02	1.50	
	9	2	0	2										0			
	Ва	0.8	0.7	0.8	0.13	0.25	1.47	0.22	0.22	3.75	1.38	1.16	0.39	4.5	8.50	4.50	
	0	3	1	4										0			
	Na	0.0	0.7	0.0	0.10	0.18	0.04	0.05	0.05	0.11	1.39	0.49	0.12	0.1	0.13	0.02	
	₂ 0	2	8	3										7			
	K_2	0.0	0.2	0.0	0.02	0.20	0.15	0.03	0.03	0.34	0.15	2.45	0.03	0.2	0.05	0.22	
	0	2	3	2										2]
	P_2	0.0	0.0	0.0	0.01	0.01	0.02	0.02	0.01	0.06	0.11	0.01	0.25	0.1	0.07	0.11]

1596 from the HZKS-1 drill core at Maoniuping.

	O ₅	2	2	2										0			
	SO	0.6	0.5	0.6	0.15	0.73	0.64	0.73	0.73	3.91	1.50	0.90	0.47	2.5	5.03	2.58	
	3	8	5	7										8			
	F	0.1	0.8	0.0	0.11	0.30	1.30	0.10	0.10	7.10	0.10	0.10	0.20	###	9.40	####	
		2	0	0										##		#	
	Cl	0.0	0.0	0.0	0.00	0.00	0.00	0.01	0.01	0.01	0.00	0.00	0.01	0.0	0.00	0.01	
		0	1	0										1			\mathbf{A}
	LOI	###	###	###	###	###	###	###	###	####	###	####	###	###	###	####	
		##	##	##	##	##	##	#	##	#	##	#	#	##	##	#	
	Tot	###	###	###	###	###	###	###	###	####	###	####	###	###	###	####	
	al	##	##	##	##	##	##	#	##	#	##	#	#	## (##	#	
	рр													\sum	2		
	m												\langle)		
	Li	2.4	31.	2.4	1.10	59.7	0.40	2.10	1.60	1.20	2.9	4.20	4.10	1.7	17.0	19.2	
		0	9	0										0			
	Ве	b.d.	b.d.	b.d.	b.d.	b.d.	b.d.	0.42	0.05	0.37	b.d.	b.d.	b.d.	5.2	1.41	3.98	
														1			
	Rb	1.6	75.	1.7	1.50	19.3	5.50	1.60	1.30	10.0	104	92.6	1.60	8.8	3.40	202	
		0	4	0							/			0			
	Sr	###	875	###	###	###	###	###	980	7870	980	8450	###	###	862	####	
		##	0	##	##	##	##	#	9	Y	9		#	##	5	#	
	Ва	742	673	757	117	223	###	###	208	8957	###	####	349	###	###	####	
		0	0	0	5	9	##	#	0		##	#	3	##	##	#	
	Zn	51.	123	51.	52.7	137	32,0	13.0	13.0	409	31.0	33.0	24.0	201	92.0	225	
		7		7										0			
	Sc	0.5	0.7	0.5	0.40	1.20	0.20	0.40	0.40	0.40	0.90	0.50	0.70	0.6	0.50	1.60	
		0	0	0 /) _								0			
	Y	207	124	211	239	201	104	124	129	74.0	152	87.0	203	147	356	145	
	Zr	2.0	97.	4.0	2.00	4.00	3.00	3.00	3.00	15.0	126	32.0	6.00	23.	8.00	277	
		0	<u> </u>	0										0			
	Nb	32.	<u>19</u> 7	33.	3.70	30.2	2.10	6.70	6.70	52.5	36.5	46.3	4.10	272	73.5	283	
		2	0	1													
	Sn	0.2	2.0	0.2	0.20	1.00	1.00	1.00	1.00	1.00	5.00	2.00	2.00	1.8	0.00	5.30	
		0	0	0										0			
X	La	644	980	655	548	832	980	541	756	1070	780	825	908	294	###	2840	
Y													0	0	##		
	Ce	135	140	136	131	154	147	933	147	1680	136	1455	###	394	###	4690	
		0	1	5	5	0	0		0		5		#	0	##		
	Pr	166	126	165	172	160	111	91.7	156	162	141	143.	173	352	480	463	
												0	3		1		
	Nd	622	496	628	680	559	338	302	552	529	480	478	308	102	###	1210	

														0	5	##		
	Sm		108	72.	108	125	89.6	41.5	36.4	76.3	57.6	70.3	57.6	209	88.	828	116	
				5											5			
	Eu		27.	16.	28.	33.1	22.3	13.5	7.77	17.6	12.4	16.8	13.2	45.5	18.	133	19.8	
			8	5	5										5			
	Gd		77.	44.	79.	94.4	60.7	26.8	19.6	47.0	28.1	46.8	31.8	106	41.	269	44.8	
			7	3	0										7			\mathbf{O}
	Tb		9.5	5.3	9.8	11.6	8.06	3.26	2.21	5.66	2.91	6.03	3.80	11.6	4.5	24.2	4.12	
			3	8	6										8		0	
	Dy		44.	25.	45.	53.5	40.3	15.0	10.0	27.1	12.4	29.3	17.9	45.7	20.	76.0	14.2	
			4	7	5										4			
	Но		8.4	4.8	8.7	10.5	7.54	2.91	1.79	5.06	2.19	5.49	3.29	7.67	3.7	11.6	2.33	
			8	0	0									\sim	3)		
	Er		21.	12.	22.	26.8	19.8	8.23	4.81	13.2	5.47	15.1	8.98	19.5	9.4	28.4	5.49	
			8	7	3									\searrow	4			
	Tm		3.0	1.8	3.2	3.72	2.87	1.23	0.68	1.89	0.80	2.14	1.29	2.59	1.3	3.56	0.70	
			4	1	2										5			
	Yb		18.	10.	18.	21.6	16.6	7.81	4.22	11.3	4.82	12.8	7.76	15.3	7.9	20.2	4.54	
			4	7	5						$\mathbf{}$				8			
	Lu		2.5	1.6	2.5	3.12	2.53	1.21	0.65	1.76	0.78	1.84	1.20	2.05	1.3	2.71	0.65	
			6	4	8				\checkmark						1			
	Hf		0.1	2.8	0.1	0.10	0.10	0.10	0.10	0.10	0.60	3.90	1.00	0.20	0.9	0.20	8.50	
			0	0	0										0			
	Та		0.0	1.5	0.0	0.05	0.10	0.10	0.10	0.10	0.40	0.90	0.50	0.10	1.2	0.50	2.97	
			7	0	7		$\langle \rangle$	\checkmark							0			
	Pb		127	900	126	142	550	141	###	538	1400	515	97.5	117	312	547	830	
					/			0	#	0					0			
	Th		3.6	12.	3.5	1.71	140	4.39	3.99	4.19	28.7	14.2	16.0	106	55.	674	156	
			1	7	7										8			
	U		23.	3.8	23.	2.61	3.70	4.00	5.39	5.11	3.68	4.80	3.60	4.30	303	161	251	
			5	∕₀≻	2													
	LR		291	309	294	287	320	295	###	302	3511	285	2972	###	836	###	9339	
	EE	\checkmark	8	2	9	3	3	4	#	7		3		#	4	##		
	HR		186	107	190	225	158	66.5	44.0	113	57.4	119	76.0	210	90.	436	76.8	
	EE														5			
	ΣR		310	319	313	309	336	302	###	314	3568	297	3048	###	845	###	9416	
\mathcal{I}	EE		4	8	9	8	1	0	#	0		3		#	4	##		
	LREE/	'	15.	28.	15.	12.8	20.2	44.4	43.5	26.8	61.1	23.9	39.1	146	92.	292	122	
	HREE		7	9	6										4			
	(La/S		10.	9.2	3.7	2.75	2.75	17.8	6.07	6.19	11.6	6.93	8.94	3.78	20.	40.5	20.1	
	m) _N		0	8	8										7			
	I				i	i	i		I		i	I	I	I	I	I	I	r.

(Sm/	Y	7.1	9.3	6.3	6.27	6.27	5.78	7.41	7.35	13.0	6.00	8.07	6.37	12.	54.4	12.3	
b) _N		8	8	7										1			
(La/Y	'b	71.	87.	24.	17.2	17.2	103	45.0	45.4	151	41.6	72.2	24.1	250	220	247	
) _N		8	1	1											2		
Υ/		24.	25.	24.	22.8	26.7	35.7	69.3	25.5	33.8	27.7	26.4	26.5	39.	30.8	62.0	
Но		4	8	3										4			
δE		0.9	0.8	0.9	0.93	0.92	1.23	0.89	0.90	0.94	0.89	0.94	0.93	0.9	0.86	0.84	
u		2	9	4										3			
δC		1.0	1.0	1.0	1.04	1.02	1.01	1.01	1.04	0.98	1.00	1.02	1.01	0.9	1.00	0.99	
е		0	1	1										4			~
C/CN	1F	97.	96.	97.	97.3	96.1	98.1	97.2	97.3	98.4	97.8	97.4	96.9	98.	97.8	97.9	
(%)		3	1	3										4	\sum		
													\searrow)	(cont	
																inue	
													\searrow			d)	
(co	nti	nued)														
Tak	ole	1 co	ntinu	ed						$\overline{\langle}$		F				•	-

(continued) 1597

1598 Table 1 continued

	R								$\langle \rangle$		Nordm	arkite			
	0														
	C F														
	r c	H7K5	H7KS	H7KS	HZK	НТК	H7K	Н7К	Н7К	Н7К	HZK	HZK	HZKS	H7KS	Н7К
	a	-1-18	-1-21	-1-21	S-1-	S-1-	S-1-	S-1-	S-1-	S-1-	S-1-	S-1-	-1-24	-1-24	S-1-
	m	B	Δ	в	27	28	32	37	38	39	19	20	Δ		26
	n. D	5	~		$\overline{}$	10	52	57	50		10	20	~	5	
	le				\mathcal{I}										
	N														
	0														
	w	$\overline{\ }$													
	ŧ.														
~	%														
	Si	2.51	####	####	2.34	1.29	0.93	2.21	0.15	9.02	####	####	####	####	####
	0		#	#							#	#		#	#
\bigcirc	2														
	Ti	0.20	0.21	0.20	0.08	0.21	0.01	b.d.	b.d.	b.d.	0.18	0.18	0.08	0.10	0.16
	0														
	2														
	А	0.76	1.81	1.79	0.36	4.93	0.03	0.06	<0.0	2.32	####	####	####	####	9.11
	I_2								1		#	#		#	

				-	1	1		1				1	1	-			
	0																
	3																
	т		1.04	0.84	0.85	1.50	0.84	1.77	1.06	0.30	0.63	1.68	1.13	2.78	3.17	1.38	
	F																
	е																
	2																
	0																$\Delta \gamma$
	3																\mathbf{X}
	F		0.33	0.30	0.22	0.27	0.31	0.42	0.40	0.25	0.27	0.33	0.24	0.31	0.25	0.24	
	е														(V	¥
	0													(\sim		
	М		0.10	0.41	0.42	0.42	0.41	0.46	0.72	0.69	0.60	0.03	0.03	0.02	0.02	0.12	
	n												~	\sim)		
	о																
	М		0.90	0.19	0.19	0.13	0.19	0.12	0.33	0.08	0.15	0.04	0.07	0.14	0.15	0.06	
	g																
	о																
	С		####	####	####	####	####	####	####	####	####	0.22	0.30	0.48	0.45	####	
	а		#	#	#	#	#	#		#)#					#	
	0									$\langle \rangle$							
	S		0.15	0.61	0.63	1.50	0.61	8.63	1.15	1.01	1.03	0.31	0.02	0.15	0.17	0.11	
	r								× ×								
	0								7								
	В		4.36	5.23	5.29	4.48	5.01	####	0.16	0.15	0.34	0.58	0.09	0.14	0.16	0.46	
	а					\sim		#									
	0				~												
	Ν		0.59	0.41	0.37	0.17	0.39	0.04	0.30	0.04	0.54	5.68	5.64	5.49	5.48	1.63	
	а																
	2																
	0		~~														
	К		0.21	1.79	1.78	0.22	0.83	0.03	0.07	0.01	1.61	5.57	6.16	6.23	4.10	4.87	
	2		$\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{$														
/	0	-	7														
	Р		0.05	0.09	0.09	0.10	0.09	0.02	0.01	<0.0	0.01	0.07	0.02	0.03	0.02	0.02	
	2									1							
	0																
	5										• -						
	S		2.87	3.31	3.35	2.56	3.21	####	0.33	0.23	0.38	0.31	0.05	0.27	0.61	1.27	
	0							#									
	3																
	F		####	2.00	2.10	####	####	0.10	0.20	0.10	0.20	0.00	0.10	0.00	0.00	2.10	
			#			#	#										

-																	-
	С		0.01	0.01	0.01	0.01	0.01	0.01	b.d.	b.d.	b.d.	0.02	0.03	0.00	0.00	0.01	
	I																
	L		####	####	####	####	####	####	####	####	####	1.83	1.24	0.52	0.94	7.16	
	0		#	#	#	#	#	#		#	#						
	I																*
	т		####	####	####	####	####	####	####	####	####	####	####	####	####	####	
	о		#	#	#	#	#	#		#	#	#	#		#	#	\mathcal{O} \mathcal{Y}
	t																
	al																
	p																
	p													(
	m													\sim	\bigcirc		
	Li		19.3	18.4	18.0	1.60	18.2	0.80	14.9	0.60	8.80	10.7	11.6	36.2	35.3	0.90	
	В		3.89	0.84	0.84	4.91	0.89	0.34	2.95	0.06	1.81	0.22	3.06	6.81	7.65	0.90	
	е												\searrow	P			
	R		200	157	148	8.60	152	1.80	2.90	0.40	51.6	189	239	213	210	0.15	
	b																
	S		1430	7380	7180	####	7180	####	>100	9500	9780	421	240	1310	1545	5930	
	r					#		#	00		\checkmark						
	В		####	####	####	####	####	####	####	1295	3000	5880	970	1395	1590	4120	
	а		#	#	#	#	#	#	\frown	/							
	Z		232	1080	1040	2020	1070	262	64.0	20.0	48.0	31.0	33.0	53.0	33.0	1080	
	n								·								
	S		1.60	0.50	0.50	0.60	0.50	0.10	0.90	0.20	0.40	0.30	2.20	0.40	0.40	0.40	
	c																
	Y		51.6	113	111	142	113	63.1	125	146	124	20.6	45.4	16.7	19.9	151	
	Z		266	226	219	20.0	230	16.0	4.00	<2.0	2.00	265	162	1/1	166	134	
	r		202	252	252	250	240	5.00	0.70	0	12.0	56.6	25.0	145	16.5	70.0	
	N h		282	252	252	259	248	5.00	0.70	0.20	12.9	50.0	25.9	14.5	10.5	70.6	
	u c		5 20	2 00	2 00	1 70	2 80	1.00	0.70	<0.2	0.50	1.00	7.00	1 70	1 50	1.60	
	n		5.50	3.00	3.00	1.70	2.80	1.00	0.70	0.2	0.50	1.00	7.00	1.70	1.50	1.00	
			2830	####	####	2830	#####	726	634	654	583	565	56.0	139	143	777	
~	a	\mathcal{I}	2000	#	#	2000	#	, 20	001	031	505	505	50.0	135	115		
	e		4690			3780	#####	910	####	1205	1070	588	98.7	214	221	1190	
	e		1050	#	#	3700	#	510		1205	10/0	500	50.7			1150	
	P		459	5554	4529	338	3845	84.1	130	130	114	53.8	10.2	####	20.2	124	
	r																
	N		1210	7200	7180	973	7170	278	469	471	405	158	34.7	58.0	60.3	429	
	d					-	_	_		_						-	
	S		115	433	428	86.2	430	37.0	65.6	66.5	56.0	15.1	7.29	7.60	8.60	66.5	
			-		-			_	-		-		-			_	l

																1
m																
Eu		19.6	64.5	62.8	18.3	62.1	9.33	15.5	16.2	13.8	3.50	1.21	1.63	1.97	17	
															.1	
Gd		43.1	119	114	42.2	113	22.0	39.1	41.3	34.9	7.25	6.12	4.39	5.49	44	
															.5	*
Tb		3.93	11.7	11.1	4.44	10.9	2.63	4.66	5.07	4.30	0.85	1.10	0.61	0.74	5.	
															80	\mathbf{A}
Dy		13.9	32.5	32.8	20.6	32.3	12.5	22.6	24.8	21.2	3.71	6.87	2.98	3.85	29	
															.2	
Но		2.22	4.84	4.70	3.63	4.81	2.31	4.21	4.58	3.98	0.64	1.51	0.55	0.65	5.	
															31	
Er		5.13	11.7	11.4	9.36	11.4	6.09	11.7	13.0	11.5	1.70	4.55	1.32	1.52	13	
												1			.7	
Tm		0.74	1.44	1.47	1.35	1.47	0.90	1.62	1.80	1.60	0.22	0.77	0.16	0.19	1.	
••••		0.71		1.17	1.55	1.17	0.50	1.02	1.00	1.00	0.22		0.10	0.15	84	
Vh		4.20	0.26	0.75	0 77	8.60	F 21	0.26	10.4	0.25	1.20	4.06	0.02	1.00	10	
YD		4.20	8.30	8.75	8.23	8.60	5.31	9.36	10.4	9.35	1.26	4.96	0.92	1.09	10	
															./	
Lu		0.62	1.51	1.47	1.27	1.48	0.83	1.31	1.41	1.33	0.19	0.81	0.13	0.15	1.	
															50	
Hf		8.00	6.50	6.40	0.80	6.60	0.50	0.30	<0.1	0.20	6.60	6.30	5.70	5.90	3.	
									0						70	
Та		2.92	1.09	1.09	1.13	1.09	0.20	b.d.	<0.0	0.05	1.00	2.70	0.59	0.69	0.	
									5						59	
Pb		839	825	799	3130	802	8950	207	91.2	194	329	22.2	99.5	58.0	17	
					\sim	\mathbf{N}									10	
Th		157	361	357	53.6	355	4.07	1.47	1.37	15.8	32.2	29.8	53.2	50.6	9.	
				\checkmark	Y										10	
U		249	120	112	303	115	3.37	0.52	0.12	4.52	16.6	6.83	12.0	12.1	62	
															.8	
LRE		9324	***	####	8025	####	2044	####	2543	2241	1383	208	440	455	26	
E		\square) #	#		#									03	
HRE		73.8	191	186	91 1	184	52.6	94 5	102	88.1	15.8	26.7	11 1	13.7	11	
			191	100	51.1	101	52.0	51.5	102	00.1	15.0	20.7		10.7	3	
SPE		0208	#####		0117		2007		2645	2220	1200	225	451	160	7	
ZRE		9398	##### 	<i>*****</i>	8117	<i>*****</i>	2097	*****	2045	2329	1399	235	451	408	27	
E .			#	#		#									16	
LREE/	HRE	126	518	534	88.1	524	38.9	26.5	24.9	25.4	87.4	7.80	39.8	33.2	23	
E													<u> </u>		.1	
(La/Sr	n) _N	15.3	42.7	41.6	20.5	40.5	12.3	40.5	40.5	40.5	23.4	10.3	11.4	10.3	8.	
															51	
(Sm/Y	′b) _N	29.9	56.3	53.2	11.4	54.4	7.58	54.4	54.4	54.4	13.0	8.58	8.99	8.58	5.	
															52	l
(La/Yb) _N	458	2404	2211	234	2202	1.00	####	2202	2202	305	88.8	103	88.8	47	
--------	----------------	-------	------	------	------	------	------	------	------	------	------	------	------------	------	----	----------
															.0	
Y/H		23.2	23.3	23.5	39.0	23.4	27.3	29.6	31.9	31.2	32.2	30.1	30.4	30.6	28	
o															.3	
δEu		0.85	0.87	0.87	0.92	0.86	1.00	0.93	0.94	0.95	1.02	0.55	0.86	0.87	0.	
															96	
δCe		1.00	1.05	1.24	0.94	1.33	0.89	1.01	1.00	1.01	0.82	1.00	0.99	1.00	0.	N
															93	
C/CMI	F(%)	96.8	97.9	98.0	98.5	98.1	97.4	97.5	98.5	98.1					2	
(con	itinu	ied)											<u>ر</u> C			
Tab	le 1	conti	nued													

(continued) 1599

1600 Table 1 continued

	-		Allert										allya	***	
	ock		Aikali §	granite								Altered	ankan gra	nite	
	Sa		НΖК	НΖК	HZKS	HZKS	HZKS	HZKS	HZKS	HZKS	нгкѕ	HZKS	HZKS-	HZKS	HZKS
	mpl		S-1-8	S-1-9	-1-14	-1-16	-1-23	-1-29	-1-30	-1-34	-1-35	-1-15	1-23A	-1-33	-1-36
	e										\rightarrow				
	No.								Á						
	wt.									\) /					
	%								$\langle \rangle$						
	SiO ₂		####	####	####	####	####	####	####	####	####	####	#####	####	####
			#	#	#	#	#	#	×	#	#	#			#
	TiO ₂		0.11	0.10	0.11	0.16	0.12	0.17	0.17	0.74	0.16	0.21	0.16	0.71	0.35
	Al_2		9.63	####	9.63	####	*****	####	####	7.57	9.19	####	9.19	9.53	####
	O ₃			#	~	#	#	#				#			#
	TFe		4.18	3.17	4.18	1.76	1.40	2.62	2.62	4.55	1.37	0.94	1.37	3.57	4.39
	₂ O ₃														
	FeO		0.25	0.41	0.41	0.42	0.33	0.26	0.24	0.32	0.31	0.31	0.22	0.25	0.29
	Mn		0.11	0.02	0.11	0.01	0.02	0.05	0.05	0.12	0.11	0.28	0.11	0.11	0.10
	0			Y											
	Mg		0.16	0.15	0.16	0.15	0.17	0.24	0.24	0.05	0.05	0.09	0.05	0.05	0.05
/	0	5	X												
	CaO		3.38	0.45	3.38	0.55	0.61	3.48	3.42	0.43	3.85	####	#####	0.32	0.13
												#			
/	SrO		0.43	0.17	0.43	0.02	0.47	0.06	0.06	0.01	0.11	1.75	0.11	0.01	0.00
	BaO		0.55	0.16	0.55	0.08	0.17	0.52	0.52	0.04	0.46	3.94	0.46	0.03	0.01
	Na_2		0.30	2.48	0.30	3.65	5.07	2.97	2.94	2.59	1.63	0.69	1.63	3.44	3.57
	0														
	K ₂ O		5.67	4.10	5.67	4.90	4.60	5.06	5.04	2.84	4.85	8.24	5.85	3.34	4.30
	P_2O		0.09	0.02	0.09	0.02	0.03	0.02	0.02	0.04	0.02	0.07	0.02	0.02	0.02

				1												1
	5															
	SO ₃		1.40	0.61	1.40	0.05	0.47	1.60	1.60	0.08	1.26	4.97	1.26	0.03	0.03	
	F		0.80	0.00	0.80	0.00	0.20	0.80	0.70	0.00	3.10	0.00	3.10	0.00	0.00	
	Cl		0.01	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	
	LOI		1.40	0.94	1.40	0.51	1.24	2.81	3.04	2.24	6.57	####	6.57	2.12	1.76	>
												#				
	Tot		####	####	####	####	####	####	####	####	####	####	#####	####	####	\mathbf{O} \mathbf{Y}
	al		#	#	#	#	#	#		#	#	#			#	
	рр														Ś	
	m															
	Li		2.10	35.3	2.10	16.8	32.4	13.8	13.9	30.4	4.10	2.70	4.10	22.3	51.6	
	Ве		1.98	7.65	1.98	2.77	0.61	4.03	3.89	307	1.78	1.73	1.78	265	253	
	Rb		256	210	256	0.00	225	250	257	990	0.03	0.25	0.03	964	1225	
	Sr		3620	1545	3620	122	1620	459	457	115	1020	####	1020	109	33.4	
												#	$\boldsymbol{\Sigma}$			
	Ва		4926	1590	4926	876	1690	4910	####	410	4380	####	4380	211	125	
										-		#				
	Zn		1740	33.0	1740	38	67.0	187	194	2950	57	72	57	1790	2790	
	Sc		1.20	0.40	1.20	1.60	0.80	1.60	1.50	0.10	0.50	0.20	0.50	0.10	0.20	
	Y		16.4	19.9	16.4	22.8	18.2	42.7	41.6	1590	47.3	57.0	47.3	1995	1100	
	Zr		175	166	175	124	154	272	266	####	380	127	380	####	3100	
									, Y	#						
	Nb		58.1	16.5	58.1	37.6	26.0	58.4	57.3	1495	125	245	125	1925	1615	
	Sn		3.30	1.50	3.30	2.70	2.00	4.40	4.80	325	3.80	3.60	3.80	299	245	
	La		757	143	757	185	122	185	170	343	278	561	278	452	367	
	Ce		1175	221	1175	269	208	271	254	1035	428	822	428	1315	1030	
	Pr		107	20.2	107	25.4	18.8	26.0	25.0	133	43.5	82.2	43.5	196	125	
	Nd		287	60.3	287	74.6	58.6	84.8	80.6	496	143	264	143	706	423	
	Sm		23.0	8.60	23.0	8.73	8.49	12.9	12.9	179	20.4	34.3	20.4	254	137	
	Eu		5.26	1.97	5.26	1.45	1.91	2.31	2.29	1.63	4.07	9.81	4.07	2.31	1.31	
	Gd	1	10.0	5.49	10.0	5.69	5.26	8.16	8.27	211	11.6	21.9	11.6	279	163	
	тр		1.02	0.74	1.02	0.80	0.73	1.24	1.24	44.9	1.37	2.54	1.37	59.5	35.3	
/	Dy	5	3.65	3.85	3.65	4.15	3.48	7.18	7.05	293	7.32	11.9	7.32	390	224	
	Но		0.65	0.65	0.65	0.79	0.61	1.53	1.43	63.0	1.37	2.31	1.37	82.5	46.2	
	Er		1.67	1.52	1.67	2.24	1.41	4.43	4.28	181	3.91	5.98	3.91	232	130	
~~	Tm		0.25	0.19	0.25	0.33	0.19	0.67	0.63	28.2	0.60	0.87	0.60	34.9	18.2	
	Yb		1.68	1.09	1.68	2.52	0.99	4.27	4.28	169	4.02	5.63	4.02	202	102	
	Lu		0.25	0.15	0.25	0.41	0.15	0.70	0.68	24.9	0.66	0.82	0.66	29.9	13.8	
	Hf		5.20	5.90	5.20	4.30	4.90	8.70	8.60	704	13.9	3.20	13.9	642	80.6	
	Та		0.38	0.69	0.38	2.12	0.90	1.72	1.75	110	2.69	1.26	2.69	122	71.8	
	Ph		5670	58.0	5670	97.0	159	220	219	994	198	239	198	327	416	
			2270	23.0	20/0	2.10		0	-13	551	190		190	517	.10	

Th		35.7	50.6	35.7	22.2	52.9	17.4	16.3	746	26.3	20.3	26.3	506	428	
U		9.26	12.1	9.26	5.56	11.5	28.4	27.9	201	102	177	102	207	99.8	
LRE		2354	455	2354	564	418	582	545	2187	916	1773	916	2925	2083	
E															
HRE		19.2	13.7	19.2	16.9	12.8	28.2	27.9	1015	30.9	51.9	30.9	1310	732	
E															
ΣRE		2373	468	2373	581	431	610	573	3202	947	1825	947	4235	2815	
E															
LREE/	Η	123	33.2	123	33.3	32.6	20.6	19.6	2.15	29.7	34.2	29.7	2.23	2.85	
REE															
(La/Sn	n)	20.6	10.3	20.6	13.2	8.97	8.93	8.26	1.20	8.51	10.2	8.51	1.11	1.67	
N													$\overline{\mathbf{i}}$		
(Sm/Y	b	14.9	8.58	14.9	3.77	9.33	3.29	3.27	1.15	5.52	6.63	5.52	1.37	1.46	
) _N															
(La/Yb)	306	88.8	306	49.9	83.7	29.4	27.0	1.38	47.0	67.7	47.0	1.52	2.44	
N															
Y/H		25.2	30.6	25.2	28.9	29.8	27.9	29.1	25.2	34.5	24.7	34.5	24.2	23.8	
о										Y					
δEu		1.06	0.87	1.06	0.63	0.87	0.69	0.68	0.03	0.81	1.09	0.81	0.03	0.03	
δCe		1.00	1.00	1.00	0.95	1.05	0.95	0.94	1.17	0.94	0.93	0.94	1.07	1.16	
C/CM	F (%))						$\mathbf{\hat{\mathbf{N}}}$							

1601 b.d. = below detection limit; C/CMF = CaO/(CaO+MgO+FeO+MnO) (in molar

1602 proportions)

Sample	Generations	δ ¹³ C _{V-PDB} ‰	δ ¹⁸ Ο _{V-SMOW} ‰
No.			
HZKS-1-6	Cal-I	-7.0	7.5
HZKS-1-7	Cal-I	-7.0	7.8
HZKS-1-17	Cal-I	-7.1	6.7
HZKS-1-22	Cal-I	-7.1	5.7
HZKS-1-2B	Cal-I	-6.6	5.4
HZKS-1-5A	Cal-II	-7.3	5.7
HZKS-1-5B	Cal-II	-6.9	7.4
HZKS-1-10A	Cal-II	-7.0	6.6
HZKS-1-2A	Cal-II	-6.8	8.4
HZKS-1-4	Cal-III	-6.7	9.2
HZKS-1-13	Cal-III	-6.9	8.4
HZKS-1-13T	Cal-III	-7.1	8.8
HZKS-1-18	Cal-III	-6.8	8.3
HZKS-1-27	Cal-III	-6.2	11.0
HZKS-1-32	Cal-III	-6.5	8.1
HZKS-1-32T	Cal-III	-6.7	9.4
HZKS-1-38	Cal-III	-6.7	10.0
HZKS-1-39	Cal-III	-6.5	9.6
HZKS-1-10B	Cal-IV	-6.9	8.2
HZKS-1-15	Cal-IV	-7.0	7.7
HZKS-1-15T	Cal-IV	-7.1	8.7
HZKS-1-26	Cal-IV	-6.8	7.9
HZKS-1-26T	Cal-IV	-6.6	7.4
HZKS-1-31	Cal-IV	-6.6	8.7

contact zone from the HZKS-1 drill core at Maoniuping.



Sampl	Miner	REE	Rb	Sr(pp	Rb/Sr	⁸⁴ Sr/ ⁸⁶	2σ	⁸⁷ Sr/ ⁸⁶	2σ	⁸⁸ S
e/spot	al	(ppm	(pp	m)		Sr		Sr		r
)	m)							(V)
HZKS-1-6-1	Cal-I	1051	0.03	7246	######	#####	#####	0.7060	#####	6.0
					#	#	#	5	#	0
HZKS-1-6-2	Cal-I	1013	0.02	7171	######	#####	#####	0.7061	#####	6.0
					#	#	#	5	Ħ	
HZKS-1-6-3	Cal-I	840	0.01	7757	######	#####	#####	0.7061	#####	5.9
					#	#	#	1	#	
HZKS-1-6-4	Cal-I	808	0.02	6422	######	#####	#####	0.7061	#####	7.9
					#	#	#	1	#	
HZKS-1-6-5	Cal-I	948	0.06	7991	######	#####	#####	0.7061	#####	6.6
					#	#	#	1	#	
HZKS-1-6-6	Cal-I	949	0.02	9325	######	#####	#####	0.7061	#####	6.2
					#	#	#	1	#	
HZKS-1-7-1	Cal-I	1135	0.06	10617	#######	#####	#####	0.7061	#####	6.8
					#	#	#	0	#	
HZKS-1-7-2	Cal-I	881	0.03	9320	######	#####	#####	0.7060	#####	7.2
					#	#	#	2	#	
HZKS-1-7-3	Cal-I	460	0.06	6356	######	#####	#####	0.7061	#####	6.7
					#	#	#	5	#	
HZKS-1-7-5	Cal-I	448	0.13	6031	######	#####	#####	0.7061	#####	5.9
		\mathbf{A}	$\langle \times \rangle$		#	#	#	1	#	
HZKS-1-17-	Cal-I	994	0.13	6368	######	#####	#####	0.7061	#####	6.5
1					#	#	#	6	#	
HZKS-1-17-	Cal-I	828	0.01	5654	######	#####	#####	0.7061	#####	5.1
2	\searrow				#	#	#	4	#	
HZKS-1-17-	Cal-I	994	0.13	6368	######	#####	#####	0.7060	#####	7.8
з у					#	#	#	3	#	
HZKS-1-17-	Cal-I	612	0.10	7432	######	#####	#####	0.7060	#####	3.7
4					#	#	#	1	#	
HZKS-1-17-	Cal-I	751	0.10	8155	######	#####	#####	0.7061	#####	5.8
5					#	#	#	9	#	
HZKS-1-17-	Cal-I	800	0.05	6105	######	#####	#####	0.7061	#####	5.0
6					#	#	#	3	#	
HZKS-1-17-	Cal-I	1082	0.10	9126	######	#####	#####	0.7062	#####	8.1
7					#	#	#	3	#	
	i					1		i		

1619 zone from the HZKS-1 drill core at Maoniuping.

	HZKS-1-17-	Cal-I	989	0.03	6909	######	#####	#####	0.7061	#####	5.4
	8					#	#	#	0	#	
	HZKS-1-17-	Cal-I	858	0.02	6295	######	#####	#####	0.7061	#####	8.8
	9					#	#	#	9	#	
	HZKS-1-22-	Cal-I	1052	0.06	7039	######	#####	#####	0.7061	#####	6.4
	1					#	#	#	7	#	
	HZKS-1-22-	Cal-I	850	0.02	6462	######	#####	#####	0.7060	#####	5.9
	3					#	#	#	6	#	
	HZKS-1-22-	Cal-I	791	0.02	6319	######	#####	#####	0.7061	#####	6.7
	4					#	#	#	0	#	
	HZKS-1-22-	Cal-I	904	0.00	6167	######	#####	#####	0.7061	#####	9.4
	7					#	#	#	9	#	
	H7KS-1-22-	Cal-I	1135	0.02	7864	#######	#####	#####	0 7061	######	64
	8	Curr	1155	0.02	/004	#	#	#	5	#	0.4
		Cal-I	1136	0.02	7486	 #######	" #####	" ######	0,7060	 ######	95
	0	Calfi	1130	0.02	7400	#	#	#	7	#	5.5
		Call	604	0.05	6169	# #######	# 	# #	/	# ######	0.1
	HZKS-1-22-	Cal-I	694	0.05	6168	###### 	#####	***	0.7061	##### 	8.1
	11					#	#	´#	1	#	
	HZKS-1-2B-	Cal-II	1441	b.d.	8659	b.d.	#####	#####	0.7061	#####	7.0
	1						#	#	5	#	
	- H7KS_1_2B_	Cal-II	1517	0.12	10935		<i>"</i>	<i></i> <i>######</i>	0 7060		59
	2 2	Cal-II	1317	0.12	10333	#	#	#	7	#	5.5
		Calli	1566	0.01	11150	л . ++++++++++	# #####	# ######	,	# ######	47
	ПZКЭ-1-2В-	CdI-II	1200	0.01		####### #	##### #	ннннн 1	0.7063	###### #	4.7
	3	<u> </u>	004 -		0520	#	#	#	0	#	5.6
	HZKS-1-2B-	Cal-II	884	D.d.	8530	b.d.	#####	#####	0.7062	#####	5.6
	4						#	#	3	#	
	HZKS-1-2B-	Cal-II	1725	0.08	8519	######	#####	#####	0.7063	#####	5.2
	7					#	#	#	1	#	
	HZKS-1-2B-	Cal-IL	1271	0.02	7500	######	#####	#####	0.7061	#####	6.5
	8	\checkmark				#	#	#	5	#	
	HZKS-1-2B-	Cal-II	1391	b.d.	7932	b.d.	#####	#####	0.7061	#####	7.6
	9						#	#	7	#	
	HZKS-1-2B-	Cal-II	1361	0.02	7897	######	#####	#####	0.7061	#####	6.0
	10					#	#	#	7	#	
X	HZKS-1-5A-	Cal-II	1804	0.11	11041	######	#####	#####	0.7061	#####	5.4
	1					#	#	#	5	#	
	HZKS-1-5A-	Cal-II	2145	0.02	12124	######	#####	#####	0.7061	#####	6.0
	2		_			#	#	#	8	#	
	H7KS-1-54-	Cal-II	1765	0.03	9869	######	#####	#####	0.7061	#####	60
	3		1,00	0.00	2005	#	#	#	4	#	
		C-1 !!	1262	0.10	8553			" #####	0 7061	" #####	50
	112K3-1-2A-	Cai-II	1202	0.10	0000	#######	#####	#####	0.7001	#####	0.0

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4					#	#	#	6	#		
HZKS-1-5A-	Cal-II	1276	0.01	10337	######	#####	#####	0.7061	#####	6.0	
5					#	#	#	3	#		
HZKS-1-5A-	Cal-II	1375	0.27	9576	######	#####	#####	0.7061	#####	5.1	
7					#	#	#	6	#		
HZKS-1-5A-	Cal-II	943	0.10	7016	######	#####	#####	0.7061	#####	7.2	1
8					#	#	#	2	#		N
HZKS-1-5A-	Cal-II	985	0.24	7359	######	#####	#####	0.7061	#####	7.9	Ň
9					#	#	#	4	#	0	
HZKS-1-5A-	Cal-II	1528	0.05	8885	######	#####	#####	0.7060	#####	8.6	
10					#	#	#	8	#		
HZKS-1-5B-	Cal-II	1503	0.18	10003	######	#####	#####	0.7061	#####	6.7	
1					#	#	#	3	#		
HZKS-1-5B-	Cal-II	998	0.01	7085	######	#####	#####	0.7063	#####	8.2	
2					#	#	#	7	#		
HZKS-1-5B-	Cal-II	969	0.27	8062	######	#####	#####	0.7062	#####	7.9	
3					#	#	#	0	#		
HZKS-1-5B-	Cal-II	747	0.05	8075	######	#####	¥####	0.7061	#####	7.8	
4					#	#	#	8	#		
HZKS-1-5B-	Cal-II	1956	0.13	17728	#######	#####	#####	0.7061	#####	8.1	
6					#	#	#	0	#		
HZKS-1-10A	Cal-II	1439	b.d.	8228	b.d.	#####	#####	0.7063	#####	6.3	
-1				$\boldsymbol{\langle}$	7	#	#	6	#		
HZKS-1-10A	Cal-II	1001	0.07	7191	######	#####	#####	0.7061	#####	5.2	
-2					#	#	#	7	#		
HZKS-1-10A	Cal-II	1288	0.04	8817	######	#####	#####	0.7061	#####	6.5	
-3		$\boldsymbol{\wedge}$			#	#	#	7	#		
HZKS-1-10A	Cal-II	1183	0.02	8407	######	#####	#####	0.7060	#####	5.4	
-4					#	#	#	5	#		
HZKS-1-10A	Cal-II	1414	0.11	8088	######	#####	#####	0.7061	#####	4.8	
-5	$\searrow^{\mathcal{V}}$				#	#	#	5	#		
HZKS-1-10A	Cal-II	1589	b.d.	8147	b.d.	#####	#####	0.7061	#####	5.6	
-6						#	#	2	#		
HZKS-1-10A	Cal-II	865	0.07	6777	######	#####	#####	0.7061	#####	4.9	
-1					#	#	#	8	#		

(continued)

1621 Table 3 continued

Sampl	Miner	REE	Rb	Sr(pp	Rb/Sr	⁸⁴ Sr/ ⁸⁶	2σ	⁸⁷ Sr/ ⁸⁶	2σ	⁸⁸ Sr
e/spot	al	(ppm)	(pp	m)		Sr		Sr		(V)

				m)								
	HZKS-1-4-1	Cal-III	2090	0.11	14195	######	#####	#####	0.7064	#####	6.4	
						#	#	#	8	#		
	HZKS-1-4-2	Cal-III	1768	0.10	13802	######	#####	#####	0.7061	#####	6.0	
						#	#	#	1	#		2
	HZKS-1-4-3	Cal-III	2105	0.48	13391	######	#####	#####	0.7064	#####	5.0	
						#	#	#	8	#		\mathbf{x}
	HZKS-1-4-4	Cal-III	1810	0.43	13958	######	#####	#####	0.7066	#####	6.2	
						#	#	#	1	#		
	HZKS-1-4-5	Cal-III	1563	0.08	8899	######	#####	#####	0.7064	#####	5.9	
						#	#	#	1	#		
	HZKS-1-13-	Cal-III	1405	0.06	9823	######	#####	#####	0.7060	****	6.1	
	1					#	#	#	2	#		
	HZKS-1-13-	Cal-III	1600	0.01	10376	######	#####	#####	0.7062	#####	5.2	
	2					#	#	#	•	#		
	HZKS-1-13-	Cal-III	1485	0.02	9012	######	#####	#####	0.7059	#####	9.8	
	4					#	#	#	2	#		
	HZKS-1-13-	Cal-III	1589	0.07	10188	######	#####	#####	0.7064	#####	9.5	
	5					#	#	#	7	#		
	HZKS-18-3	Cal-III	1379	0.01	7158	######	#####	#####	0.7063	#####	6.0	
						#	#	#	4	#		
	HZKS-18-4	Cal-III	2201	0.26	10909	#######	#####	#####	0.7063	#####	6.0	
					\sim	#	#	#	6	#		
	HZKS-18-5	Cal-III	1487	0.01	8103	######	#####	#####	0.7062	#####	6.9	
						#	#	#	8	#		
	HZKS-1-32-	Cal-III	2135	0.05	11430	######	#####	#####	0.7062	#####	3.2	
	1		\sim			#	#	#	2	#		
	HZKS-1-32-	Cal-III	1704	b.d.	10571	b.d.	#####	#####	0.7062	#####	3.2	
	3						#	#	7	#		
	HZKS-1-32-	Cal-III	1358	0.03	8910	######	#####	#####	0.7063	#####	3.2	
	4					#	#	#	8	#		
	HZKS-1-38-	Cal-III	1478	0.03	11669	######	#####	#####	0.7061	#####	6.5	
	1) '					#	#	#	8	#		
	HZKS-1-38-	Cal-III	1826	0.50	11169	######	#####	#####	0.7067	#####	5.5	
X	2					#	#	#	0	#		
	HZKS-1-38-	Cal-III	1805	0.02	7718	######	#####	#####	0.7061	#####	5.8	
	3					#	#	#	7	#		
	HZKS-1-38-	Cal-III	1808	0.26	10088	######	#####	#####	0.7061	#####	6.9	
	4					#	#	#	3	#		
	HZKS-1-38-	Cal-III	1507	0.10	9085	######	#####	#####	0.7062	#####	2.4	
	5					#	#	#	6	#		

	HZKS-1-39-	Cal-III	1996	0.19	10672	######	#####	#####	0.7061	#####	1.5	
	1					#	#	#	8	#		
	HZKS-1-39-	Cal-III	1894	b.d.	14159	b.d.	#####	#####	0.7062	#####	4.9	
	2						#	#	9	#		
	HZKS-1-39-	Cal-III	2249	0.13	14931	######	#####	#####	0.7060	#####	6.0	
	3					#	#	#	4	#		
												\mathbf{A}
	HZKS-1-2A-	Cal-IV	1082	0.50	9126	######	#####	#####	0.7062	#####	6.5	
	1					#	#	#	6	#		
	HZKS-1-2A-	Cal-IV	1460	0.88	9851	######	#####	#####	0.7061	#####	6.8	
	2					#	#	#	0	#	ノ	
	HZKS-1-2A-	Cal-IV	1549	0.05	6105	######	#####	#####	0.7062	#####	3.9	
	3					#	#	#	8	#		
	HZKS-1-2A-	Cal-IV	895	0.00	6264	######	#####	#####	0.7062	#####	4.6	
	4					#	#	#	0	#		
	HZKS-1-2A-	Cal-IV	845	0.00	7698	######	#####	#####	0.7063	#####	4.2	
	5					#	#	#	5	#		
	HZKS-1-10B	Cal-IV	1166	0.01	9568	######	#####	#####	0.7062	#####	5.6	
	-1					#	#	#	6	#		
	HZKS-1-10B	Cal-IV	1210	0.05	9404	######	#####	#####	0.7063	#####	6.0	
	-2					#	#	#	6	#		
	HZKS-1-10B	Cal-IV	1436	0.00	8639	######	#####	#####	0.7065	#####	5.9	
	-3					#	#	#	3	#		
	HZKS-1-10B	Cal-IV	1416	0.02	10705	######	#####	#####	0.7061	#####	8.6	
	-4					#	#	#	1	#		
	HZKS-1-10B	Cal-IV	818	0.06	6182	######	#####	#####	0.7065	#####	3.0	
	-5			$\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{\mathbf{$		#	#	#	5	#		
	HZKS-1-10B	Cal-IV	845	0.00	6344	######	#####	#####	0.7063	#####	5.1	
	-6					#	#	#	0	#		
	HZKS-1-15-	Cal-IV	1851	0.85	10216	######	#####	#####	0.7063	#####	5.1	
	1	Z Y İ				#	#	#	3	#		
	HZKS-1-15-	Cal-IV	1334	0.03	7160	######	#####	#####	0.7062	#####	7.0	
	2					#	#	#	0	#		
	HZKS-1-15-	Cal-IV	1744	0.03	5545	######	#####	#####	0.7062	#####	5.4	
$\mathbf{\Omega}$	3					#	#	#	1	#		
	HZKS-1-15-	Cal-IV	1627	0.03	8096	######	#####	#####	0.7064	#####	5.1	
	7					#	#	#	5	#		
	HZKS-1-15-	Cal-IV	1288	0.01	7485	######	#####	#####	0.7064	#####	5.8	
	8					#	#	#	6	#		
	HZKS-1-26-	Cal-IV	1242	0.77	10399	######	#####	#####	0.7060	#####	6.1	
	1					#	#	#	9	#		
	HZKS-1-26-	Cal-IV	1778	0.08	12348	######	#####	#####	0.7059	#####	6.2	
												ł

2					#	#	#	4	#		
HZKS-1-26-	Cal-IV	1727	0.07	12029	######	#####	#####	0.7060	#####	8.3	
3					#	#	#	6	#		
HZKS-1-26-	Cal-IV	1247	0.01	10419	######	#####	#####	0.7058	#####	8.1	
4					#	#	#	5	#		
HZKS-1-26-	Cal-IV	1413	0.05	8806	######	#####	#####	0.7060	#####	5.9	
5					#	#	#	3	#		
HZKS-1-31-	Cal-IV	1591	0.03	14190	######	#####	#####	0.7060	#####	8.9	
1					#	#	#	5	#		
HZKS-1-31-	Cal-IV	827	0.23	8403	######	#####	#####	0.7060	#####	7.8	
2					#	#	#	7	#		
HZKS-1-31-	Cal-IV	1352	0.48	13035	######	#####	#####	0.7060	#####	7.3	
3					#	#	#	9	#		
HZKS-1-31-	Cal-IV	2775	0.63	14242	######	#####	#####	0.7063	#####	7.7	
4					#	#	#	0	#		
HZKS-1-31-	Cal-IV	1489	0.35	7244	######	#####	#####	0.7061	#####	6.5	
5					#	#	#	8	#		
HZKS-1-31-	Cal-IV	1956	0.50	7432	######	#####	×######	0.7060	#####	7.0	
6					#	#	#	4	#		
HZKS-1-31-	Cal-IV	1012	0.19	10858	#######	#####	#####	0.7060	#####	7.1	
7					#	#	#	9	#		
HZKS-1-31-	Cal-IV	1126	0.02	7282	######	#####	#####	0.7061	#####	6.5	
8				$\boldsymbol{<}$)	#	#	#	5	#		
HZKS-1-31-	Cal-IV	1385	0.10	8155	######	#####	#####	0.7059	#####	5.6	
9					#	#	#	7	#		
HZKS-1-31-	Cal-IV	1079	0.13	6929	######	#####	#####	0.7066	#####	6.2	
10			Y		#	#	#	0	#		
HZKS-1-31-	Cal-IV	2099	0.05	14643	######	#####	#####	0.7063	#####	6.8	
11					#	#	#	0	#		
HZKS-1-31-	Cal-IV	2576	0.06	15274	######	#####	#####	0.7062	#####	6.8	
12	$\sum Y$				#	#	#	5	#		

1622 b.d. = below detection limit

1623

1625

Sample/	Mineral	¹³ C/ ¹² C ture	σ	С	δ ¹³ C _{PBD} (‰)
spot				intensity	
				(V)	
HZKS-1-6-1	Cal-I	###########		14.34	-6.70
			0.0000016		
HZKS-1-6-2	Cal-I	###########		14.50	-6.01
			0.0000015		
HZKS-1-6-3	Cal-I	###########		14.52	-6.28
			0.0000025		X
HZKS-1-6-4	Cal-I	###########		14.31	-5.46
			0.000033		
HZKS-1-6-5	Cal-I	###########		14.17	-3.28
			0.0000015		
HZKS-1-6-6	Cal-I	###########		14.28	-4.74
			0.0000033		
HZKS-1-7-1	Cal-I	###########		15,83	-5.61
			0.0000012	×	
HZKS-1-7-2	Cal-I	###########		14.68	-5.73
			0.0000021		
HZKS-1-7-3	Cal-I	#########	× ×	15.67	-6.55
			0.0000014		
HZKS-1-7-5	Cal-I	###########		15.98	-6.05
			0.0000014	10.07	
HZKS-1-17-1	Cal-I	###########	0.000001.6	16.87	-5.14
			0.000016	16.01	F (1
HZK5-1-1/-2	Car-i	#######################################	0.000012	10.81	-5.01
	Cal.1	#######################################	0.000013	16.40	E CO
HZK3-1-17-5	Cal-I	*****	0.000015	10.42	-5.09
1746 1 17 4	Call		0.0000015	16.24	Г 20
HZN3-1-17-4	Cal-I	*****	0.000015	10.34	-5.30
H7KS_1_17_5	Cale	#######################################	0.0000015	16.39	_/ 51
112NJ-1-1/-J	Cal-i	~~~~	0 000013	10.30	-4.71
H7KS_1_17_6	Cal-I	###########	0.0000013	16 55	-4 07
11ZNJ-1-17-0	Cal-i	~~~~	0 0000011	10.33	-4.07
H7KS-1-17-7	Cal-I	##########	0.000011	16 18	-5.64
			0.0000012	10.10	5.04
	1	1	0.000012	1	

1627 contact zone from the HZKS-1 drill core at Maoniuping.

							1
	HZKS-1-17-8	Cal-I	###########		16.16	-4.65	
				0.0000012			
	HZKS-1-17-9	Cal-I	##########		11.87	-6.55	
				0.0000018			
	HZKS-1-22-1	Cal-I	###########		14.78	-6.21	
				0.0000018			
	HZKS-1-22-3	Cal-I	###########		14.84	-5.76	
				0.0000015			
	H7KS-1-22-4	Cal-I	############		14.60	-5.36	\mathbf{o}
		0011		0 0000020			
	ד ככ 1 איקע	Call	#######################################	0.0000020	14 59	6.01	
	HZK3-1-22-7	Cal-i	*****	0.000015	14.50	-0.01	
				0.0000015	44.54	574	
	HZKS-1-22-8	Cal-I	##########		14.51	-5.74	
				0.0000014			
	HZKS-1-22-9	Cal-I	###########		14.73	-6.16	
				0.0000019			
	HZKS-1-22-11	Cal-I	###########		14.45	-5.36	
				0.0000019		Y	
					\sim		
	HZKS-1-2B-1	Cal-II	###########	×	14.23	-6.87	
				0.0000026			
	HZKS-1-2B-2	Cal-II	###########		14.25	-5.71	
				0.0000022			
	HZKS-1-2B-3	Cal-II	##########		14.34	-6.10	
				0.0000022			
	HZKS-1-2B-4	Cal-II	##########	r	13.46	-6.14	
	-		\searrow	0 0000020		_	
	H7KS-1-2B-7	Cal-II		0.000020	14 13	-7 19	
				0 0000022	14.15	7.15	
				0.0000022	12 50	6.62	
	HZK3-1-2B-8	Calein	*****	0.000001.0	13.50	-0.03	
				0.0000018			
	HZKS-1-2B-9	Cal-II	###########		13.94	-6.50	
,				0.0000023			
	HZKS-1-2B-10	Cal-II	###########		14.05	-7.16	
	X			0.0000025			
	HZKS-1-5A-1	Cal-II	###########		14.27	-7.91	
				0.0000037			
	HZKS-1-5A-2	Cal-II	###########		14.28	-4.67	
				0.0000014			
	HZKS-1-5A-3	Cal-II	###########		14.35	-5.20	
				0.0000013			
			#######################################	0.000013	1/ 20	-5 /12	
	HZK3-1-3A-4	Cal-II	<i>***********</i>		14.29	-5.45	J

			0.0000019			
HZKS-1-5A-5	Cal-II	##########		16.21	-7.00	
			0.0000014			
HZKS-1-5B-1	Cal-II	###########		10.92	-6.28	
			0.0000017			
HZKS-1-5B-2	Cal-II	##########		10.87	-6.03	
			0.0000020			
HZKS-1-5B-3	Cal-II	##########		10.73	-6.01	
			0.0000019			
HZKS-1-5B-4	Cal-II	##########		10.75	-5.24	
			0.0000017			
HZKS-1-5B-6	Cal-II	##########		9.95	-5.73	
			0.0000025		~	
HZKS-1-10A-1	Cal-II	##########		16.87	-7.77	
			0.0000016			
HZKS-1-10A-2	Cal-II	##########		17.77	-6.58	
			0.0000018			
HZKS-1-10A-3	Cal-II	##########		16.42	-6.08	
			0.0000012			-
HZKS-1-10A-4	Cal-II	##########		16.76	-5.95	
			0.0000012			
HZKS-1-10A-5	Cal-II	##########		16.49	-6.22	
			0.0000013			-
HZKS-1-10A-6	Cal-II	#########		16.24	-5.86	
			0.0000015			

1628 (continued)

1629 Table 4 continued

Sample/	Mineral	¹³ C/ ¹² C ture	σ	С	δ ¹³ C _{PBD} (‰)
spot	Y			intensity	
				(V)	
HZKS-1-4-1	Cal-III	###########		14.35	-8.20
			0.0000031		
HZKS-1-4-2	Cal-III	###########		16.18	-7.94
			0.0000015		
HZKS-1-4-3	Cal-III	###########		16.04	-7.34
			0.0000019		
HZKS-1-4-4	Cal-III	###########		16.22	-7.16
			0.0000013		
HZKS-1-4-5	Cal-III	###########		14.29	-7.93

					
			0.0000027		
HZKS-1-13-1	Cal-III	###########		17.81	-6.92
			0.0000015		
HZKS-1-13-2	Cal-III	###########		17.93	-7.32
			0.0000016		
HZKS-1-13-4	Cal-III	############		17.39	-7.77
			0.0000014		
HZKS-1-13-5	Cal-III	###########		17.11	-6.77
			0.0000017		
H7KS-18-3	Cal-III	############		17 14	-9 97
			0 0000017		
H7KS-18-4	Cal-III	############	0.000001/	16 97	-8 09
11210 10 4	curm		0.0000014	10.57	0.05
		#######################################	0.0000014	0 07	0 14
1121(3-10-3	Cal-III	*****	0.000024	0.02	-0.14
	Caluu		0.0000024	15.90	6.01
HZK3-1-32-1	Cal-III	############	0.0000011	15.80	-6.91
			0.0000011		
HZKS-1-32-3	Cal-III	###########		15.63	-7.92
			0.0000015		
HZKS-1-32-4	Cal-III	###########	l k	15.66	-6.55
			0.0000013		
HZKS-1-38-1	Cal-III	###########		15.60	-6.92
			0.0000012		
HZKS-1-38-2	Cal-III	#########		15.90	-5.84
			0.0000014		
HZKS-1-38-3	Cal-III	#######################################		15.73	-7.08
			0.0000013		
HZKS-1-38-4	Cal-III	###########		15.59	-7.05
	\sim		0.0000014		
HZKS-1-38-5	Cal-III	###########		16.30	-6.71
	Y		0.0000013		
HZKS-1-39-1	Cal-III	###########		16.45	-6.78
			0.0000017		
HZKS-1-39-2	Cal-III	###########		16.65	-7.25
			0.0000017		
HZKS-1-39-3	Cal-III	###########		16.42	-7.46
			0 0000015		
HZKS-1-2A-1	Cal-IV	###########		9.39	-6.38
			0.0000028		
HZKS-1-2A-2	Cal-IV	###########		9.13	-6.81
			0.0000026		
	1		5.000020		1

	HZKS-1-2A-3	Cal-IV	###########		8.91	-7.67	
				0.0000022			
	HZKS-1-2A-4	Cal-IV	###########		9.55	-5.14	
				0.0000032			
	HZKS-1-10B-1	Cal-IV	###########		16.95	-7.43	
				0.0000016			
	HZKS-1-10B-2	Cal-IV	##########		15.52	-7.87	
				0.0000015			
	HZKS-1-10B-3	Cal-IV	##########		16.39	-7.44	
				0.0000017			
	HZKS-1-10B-4	Cal-IV	##########		17.06	-5.96	
				0.0000014			$\langle \mathbf{n} \rangle$
	HZKS-1-10B-5	Cal-IV	###########		16.38	-6.65 🔨	
				0.0000015			
	HZKS-1-10B-6	Cal-IV	###########		16.60	-5.87	
				0.0000015			
	HZKS-1-15-1	Cal-IV	##########		16.81	-7.17	
				0.0000016		Y	
	HZKS-1-15-2	Cal-IV	##########		16.78	-7.03	
				0.0000021			
	HZKS-1-15-3	Cal-IV	##########		17.34	-6.72	
				0.0000014	/		
	HZKS-1-26-1	Cal-IV	##########	$\langle \rangle'$	17.05	-6.18	
				0.0000016			
	HZKS-1-26-2	Cal-IV	############		16.76	-8.61	
			$\overline{\mathbf{A}}$	0.0000017			
	HZKS-1-26-3	Cal-IV 🔨	###########		16.74	-6.32	
				0.0000016			
	HZKS-1-26-4	Cal-IV	###########		16.63	-6.73	
	\checkmark			0.0000015			
	HZKS-1-31-1	Cal-IV	##########		15.75	-5.55	
				0.0000014			
/	HZKS-1-31-2	Cal-IV	##########		15.70	-5.95	
				0.0000015			
	HZKS-1-31-3	Cal-IV	##########		15.83	-6.35	
$\mathbf{\mathbf{y}}$				0.0000018			
	HZKS-1-31-4	Cal-IV	##########		15.55	-4.78	
				0.0000012			
	HZKS-1-31-5	Cal-IV	##########		15.47	-5.64	
				0.0000015			
	HZKS-1-31-6	Cal-IV	##########		15.11	-5.35	

R

				0.0000018			
	HZKS-1-31-7	Cal-IV	###########		15.12	-6.05	
				0.0000015			
	HZKS-1-31-8	Cal-IV	###########		15.64	-6.74	
				0.0000016			
	HZKS-1-31-9	Cal-IV	############		15.34	-6.28	
				0.0000014			
	H7KS-1-31-10	Cal-IV	#######################################		15.46	-5.83	
		curre		0.000015	10110	5.00	\mathbf{o}
		Cal IV		0.0000015	15.22	Γ 41	
	П2К3-1-51-11	Cal-IV	*******	0.0000047	15.55	-5.41	
		- • • • •		0.0000017			
	HZKS-1-31-12	Cal-IV	###########		14.80	-6.43	
				0.0000017		\checkmark	
1630							
1050							¢
1621							
1051							
1 (2)						\succ	
1632							
1633							
1634				\mathbf{A}			
				\mathbf{N}			
1635							
1636							
1050		A	$\langle \rangle^{\gamma}$				
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1638	\sim	Y					
1639							
1640							
	Y						
1641	-						
1642							
1042							

1643 FIGURE CAPTIONS

1644 Fig. 1. Geology and location maps of the Maoniuping deposit. (a) Regional

- 1645 tectonic map of China, showing the eastern Tibetan orogen (outlined red). (b)
- 1646 Simplified tectonic map of eastern Tibetan Orogen (modified after Yin & Harrison,
- 1647 2000; Hou and Cook, 2009), including the Mianning-Dechang REE belt (yellow
- 1648 square) and the location of the Maoniuping deposit (highlighted in red). (c) Detailed
- 1649 geological map of the Maoniuping deposit displaying an overview of the
- 1650 carbonatite-nordmarkite complexes (modified from Yuan et al., 1995). The Dagudao
- 1651 and Guangtoushan sections are highlighted by dashed blue lines.
- 1652

Fig. 2. Drill core section of HZKS-1 from 0 to 1100 m showing lithologies,
sampling positions, and the distribution of REE mineralization.

1655

Fig. 3. Representative photos of investigated samples from the HZKS-1 drill core. 1656 (a) Fresh fine-grained, gray nordmarkite. (b) Altered nordmarkite with fenitization 1657 1658 and hematization. (c) Alkaline granite containing quartz, K-feldspar and interstitial 1659 clinopyroxene and mica. (d) Contact zone between carbonatites and nordmarkite, 1660 including aegirine, arfvedsonite, pseudo-hexagonal phlogopite (Phl-II), K-feldspar as 1661 well as bastnäsite-(Ce) aggregates; (e) Reddish and altered granite with carbonation. 1662 (f) Nordmarkite crosscut by the carbonatite I. (g and h) Carbonatite I containing calcite (Cal-I) cumulates, disseminated phlogopite (Phl-I) and riebeckite. (i) 1663

1665 Nordmarkite crosscut by carbonatite II, which are predominantly calcite cumulates

1666 (Cal-III). (k) Carbonatite II, comprising centimeter-scales calcite (Cal-III), comb-like

1667 phlogopite (Phl-II), purple fluorite, quartz, baryte and bastnäsite-(Ce). Abbreviations:

- 1668 Aeg = aegirine, Agt = aegirine-augite, Arf = arfvedsonite, Brt = baryte, Bsn
- 1669 bastnäsite-(Ce), Cal = calcite, Kfs = K-feldspar, Phl = phlogopite, Py = pyrite, Qtz =
- 1670 quartz, Rbk = riebeckite.
- 1671

1672 Fig. 4. Photomicrographs, optical microscope cathodoluminescence (OP-CL) and backscattered electron (BSE) images of carbonatite I. (a) Cross-polarized light 1673 image of the fine-grained calcites (Cal-I), showing polygonal mosaic texture. (b) 1674 1675 Plane-polarized light image showing primary phlogopite (Phl-I) with riebeckite aggregates. (c) Euhedral calcites from a thick section, showing colorful extinction in 1676 cross-polarized light. (d) Plane-polarized light and BSE (yellow square) images of 1677 1678 needle-like molybdenite aggregates with inhomogeneous distribution in calcites. (e) 1679 BSE image of the calcite cumulates. (f) Euhedral riebeckite inclusions enclosed in 1680 primary phlogopite (Phl-I). (g) BSE image of carbonatite I at depth of 929 m, showing 1681 Cal-II hosting bastnäsite-(Ce) inclusions. (h) BSE image showing part of Fig. 6g 1682 where bastnäsite-(Ce) inclusions exhibit cleavage-aligned trails within Cal-II. (i) 1683 Optical CL image of homogeneous calcite grains with bright orange to blood-red

1684 luminescence and slight core-rim zoning. Abbreviations: Fl = fluorite, Mol = 1685 molybdenite; others are the same with those Fig. 3.

1686

Fig. 5. Photomicrographs of representative textures and mineral assemblages of 1687 the carbonatite II from the HZKS-1 drill core. (a) Plane-polarized image 1688 1689 coarse-grained calcites (Cal-III) that occur with elliptic purple fluorite and contain baryte along fractures. (b) Plane-polarized image of coarse-grained fluorite and 1690 porous baryte distributed with anhedral calcites, both of which are intergrown with 1691 1692 comb-like bastnäsite-(Ce) aggregates. (c) Plane-polarized image showing euhedral needle-like arfvedsonite and patchy aggirine associated with coarse-grained calcites 1693 (Cal-III). (d) Plane-polarized image of inequigranular texture of the carbonatite II, 1694 1695 showing deformation in the brittle-ductile transition regime. (e) Cross-polarized image of interstitial phlogopite crystals along amoeboid calcite grain boundaries. (f) 1696 Cross-polarized image of elongated calcite crystals extending over 2 mm in certain 1697 orientations. (g) Cross-polarized image showing a curved boundary between 1698 1699 coarse-grained calcites (Cal-III) that exhibit bent and thick-twinned lamellae. (h) BSE 1700 image of gray fluorite and calcite crosscut by late baryte. (i) BSE image of acicular 1701 arfvedsonite, Cal-III and patchy aegirine and anhedral quartz with large pores and 1702 intergrown with baryte and bastnäsite-(Ce). (j) Parisite-(Ce) and bastnäsite-(Ce) with 1703 patchy zoning under BSE image. (k) Coarse-grained calcites showing flame-like core 1704 fading to dark orange rim under OP-CL. Abbreviations as in Fig. 3; Fl = fluorite, par
1705 = parisite-(Ce).

1706

1707 Fig. 6. Representative photomicrographs (plane-, and cross-polarized light) and 1708 BSE images of the contact zones. (a) Characteristic carbonatite-wall rock contact 1709 with large proportion of silicate minerals (phlogopite, aegirine, arfvedsonite, microcline) and REE mineralization. (b) Euhedral core-rim zoned phlogopite (Phl-II) 1710 enclosed in calcite (Cal-IV) matrix. (c) Part of Fig. 6g where pseudo-hexagonal 1711 1712 aegirine is surrounded by arfvedsonite aggregates and less-porous baryte. (d) Local 1713 area of Fig. 6b showing phlogopite with dark gray to light gray core-to-rim zoning, 1714 arfvedsonite inclusions, and surrounding euhedral arfvedsonite aggregates. (e) 1715 Euhedral laths of bastnäsite-(Ce) showing planar boundary with calcites. (f) Porous baryte and patchy againing distributed along margin of bastnäsite-(Ce) crystals. (g) 1716 Phlogopite, quartz, and agairine in proximity to bastnäsite-(Ce) and surrounded by 1717 porous baryte. Abbreviations as shown in Fig. 3, and Mc = microcline. 1718 1719

1720 Fig. 7. Chondrite-normalized whole-rock REE patterns for the carbonatite I and
1721 II, nordmarkite, and alkali granite (chondrite values from McDonough & Sun,
1722 1995).

1724	Fig. 8. Major	compositions o	of the four	generations of	f calcites in car	bonatite I, II
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1725 and the contact zone as box plot diagrams. The star symbols represent outliers.

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- 1727 Fig. 9. Chemical variations of the four generations of calcites in carbonatite I, II
- 1728 and the contact zone. R^2 means the linear correlation coefficient.

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- 1730 Fig. 10. Trace element compositions of calcites from carbonatite I, II and the
- 1731 contact zone as determined by LA-ICP-MS. (a) Chondrite-normalized REE
- 1732 patterns of four calcite generations (green dashed outline is REE patterns of Cal-IV;
- 1733 chondrite values from McDonough & Sun, 1995). (b) (La/Sm)_N versus Y/Ho scatter
- 1734 diagrams (CHARAC field from Bau, 1996)

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Fig. 11. Compositional variation of micas from carbonatite I, nordmarkite and
the contact zone. (a) Binary of Mg + Si versus ^{IV}Al + ^{VI}Al (tetrahedral and octahedral
occupancy in a.p.f.u.). Dashed line representing phlogopite–eastonite substitution
trend. (b) and (c) Fe / (Fe + Mg) (molar ratios) and Al versus Ti, respectively. (d) to (f)
Ti, Na, Mn versus Mg#, respectively. (g) and (h) Na versus K and F, respectively. (i)
Mg# versus IV(F).

1743 Fig. 12. Bulk C and O isotopic compositions of calcite from carbonatite I, II, and 1744 the contact zone from the HZKS-1 drill core. (a) δ^{13} C versus δ^{18} O plots of calcite

1745 showing a comparison between Maoniuping and other carbonatites. Typical trend 1, 2 1746 and 3 for magmatic fluid-carbonate interaction, magma fluid-carbonate interaction 1747 plus minor decarbonation, and pure decarbonation, respectively (Whitley et al., 2019). (b) δ^{13} C versus δ^{18} O plots of calcite from Maoniuping carbonatites (PIC (Primary 1748 Igneous Carbonatite) and mantle fields from Taylor et al., 1967 and Deines, 1992 1749 respectively). Comparative C-O isotopic data including Mianning-Dechang belt (e.g., 1750 Maoniuping (Yang et al., 1998; Niu et al., 2002; Xu et al., 2003; Hou et al., 2006, 1751 2015), Dalucao and Lizhuang (Liu & Hou, 2017), Miaoya (Xu et al., 2014; Ying et al., 1752 1753 2020), Bayan Obo (Yang et al., 2009), and Qieganbulake (Ye et al., 2013) 1754 carbonatites.

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Fig. 13. The in-situ Sr isotopic variation of the four calcite generations at
Maoniuping. The ranges of whole-rock ⁸⁷Sr/⁸⁶Sr ratios for nordmarkite and
carbonatite from Wang et al. (2001), Xu et al. (2003), and Hou et al. (2006, 2015).

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1760Fig. 14. Modal calculations of equilibrium and fractional crystallization showing1761differentiation degree of carbonatitic magma. (a) Equilibrium crystallization1762modeling trends. (b) to (d) Fractional crystallization modeling with variable1763calcite:amphibole:fluorite ratios and differentiation increments (Δ F) of 0.01, 0.05 and17640.1. Whole-rock compositions based on samples HZKS-1-22 (orange curve),1765HZKS-1-17 (purple curve), and HZKS-1-6 (black curve). F is degree of fractional

1767 symbols represent measured data of the four calcite generations.

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1769	Fig. 15. The schematic chemical variation of trace element compositions of
1770	calcite in carbonatite I, II and the contact zone from the HZKS-1 drill core. (a)
1771	δEu versus $\delta Ce.$ (b and c) Chondrite-normalized diagrams of REE versus (La/Yb)_N
1772	and (La/Yb) _N versus Y/Ho, respectively (chondrite values from McDonough and Sun,
1773	1995). (d) Sr/Y versus Y/Y*. Trace-element compositions of calcites for comparison
1774	are selected from Oka (Chen & Simonetti, 2013), Miaoya (Ying et al., 2020), Dalucao
1775	(Liu & Hou, 2017), and Muluozhai (Zheng et al., 2021). Other calcite data of global
1776	carbonatites and typical trend for magma fractionation, hydrothermal reworking and
1777	supergene oxidation in are compilated from Chakhmouradian et al. (2016b).

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1779 Fig. 16. Monte Carlo simulations of Sr isotopic compositions and relationship to Sr and REE concentrations, indicating more radiogenic ⁸⁷Sr/⁸⁶Sr ratios in calcites 1780 from carbonatite II and contact zone requires the ~40% contamination with 1781 nodrmarkite. In modal calculation, randomly input selecting values set 4×10^8 . The 1782 initial ⁸⁷Sr/⁸⁶Sr ratios of are 0.70605 for carbonatites according to our analyses of 1783 Cal-I and 0.70686–0.70700 for nordmarkite (Xu et al., 2003; Hou et al., 2006). Based 1784 1785 on the whole-rock results of this study and previous research, Sr contents are selected from 8000 to 25000 ppm for carbonatites (Hou et al., 2015) and 500 to 3000 ppm for 1786

1787 nordmarite (Wang et al., 2001); initial REE contents are set from 1000 to 3000 ppm

1788 for carbonatites (Wang et al., 2001; Hou et al., 2006) and 200 to 800 ppm for

1789 nordmarkite (Xu et al., 2003; Hou et al., 2015).

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Fig. 17. The schematic variation of in-situ Sr-C isotopic compositions of calcites. 1791 (a) REE versus ⁸⁷Sr/⁸⁶Sr. (b) C-Sr isotopic decoupling trend from initial 1792 early-crystalizing to late calcite. (c) REE versus δ^{13} C. (d) An intergrade comparison 1793 of δ^{13} C values among four calcite generations (in-situ), calcite (bulk), bastnäsite-(Ce), 1794 1795 carbonatites at Maoniung. The results of bulk C isotopic compositions for calcite are from this study (red) and previous studies (purple; Niu et al., 2002; Xu et al., 2003; 1796 Hou et al., 2006, 2015; δ^{13} C values of bastnäsite-(Ce) and mantle values are from Liu 1797 & Hou, 2017, and Deines, 1992, respectively). (e) Chondrite-normalized $(La/Yb)_N$ 1798 versus δ^{13} C (chondrite values from McDonough & Sun, 1995). (f) Raleigh 1799 decarbonation modeling illustrating relationship between Fco₂ (the mole fraction of C 1800 1801 remaining in the melt after decarbonation, Valley, 1986) and C isotopic composition of evolved carbonatitic melt. The initial δ^{13} C values of carbonatitic melt is taken from 1802 1803 the average Cal-I value of -5.3% (pink circle).

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Fig. 18. A schematic evolutionary model of carbonatitic magma illustrating
 chemical variation and mineral assemblages during process of fractional
 crystallization and contamination of carbonatitic melt by silicate wall rock. (a) A

2021), with carbonatite melt fractional crystallization and contamination by nordmarkite at shallow crust. (b1) Incipient magmatic texture and primary mineral assemblages of carbonatite I that derived from a low degree (F = ~ 0.9) of fractional crystallization of carbonatitic melt. (b2) Ascent and further fractional crystallization (F = 0.8-0.7) of carbonatite melt cause primary bastnäsite-(Ce) crystallization along cleavage-aligned trails within Cal-II. (b3) Contamination-induced interaction between carbonatite melt and nordmarkite counterparts leads to scavenging of Na, K, and F from carbonatite melt and incorporation of Si and Al from wallrocks by formation of phlogopite, clinopyroxene, microcline, fluorite and bastnäsite-(Ce). (b4) Protracted fractional crystallization (F = 0.6-0.5) and repeated contamination result in huge REE accumulation, decarbonation-triggered carbonatite magma ascent, crystallization of super coarse-grained calcite (Cal-IV) with silicates, and textural deformation of carbonatite II.

schematic diagram show carbonatite magma ascent (modified from Walter et al.,

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Fig. 2.







Fig. 5.





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Fig. 7.











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Fig. 15.

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